

Sync rad: Radiated power

- “Circular” orbits:

$$P_\gamma = \frac{2}{3} r_e m c^3 \frac{\gamma^4 \beta^4}{\rho^2}, \quad r_e = \frac{e^2}{4\pi\epsilon_0 m c^2}.$$

Radiation in forward direction with opening angle $\propto \gamma^{-1}$

Classical radii of a few species:

$$r_e = 2.82 \times 10^{-15} \text{ m}$$

$$r_p = 1.53 \times 10^{-18} \text{ m}$$

$$r_{\text{Au}} = 4.88 \times 10^{-17} \text{ m}$$

$$r_{\text{Pb}} = 4.98 \times 10^{-17} \text{ m}$$

$$P_\gamma \propto \left(\frac{q}{m}\right)^4 U^2 B^2, \quad \text{since } \rho = \frac{p}{qB_\perp} \text{ for fixed radius.}$$



Energy loss per turn

- Energy loss per turn:

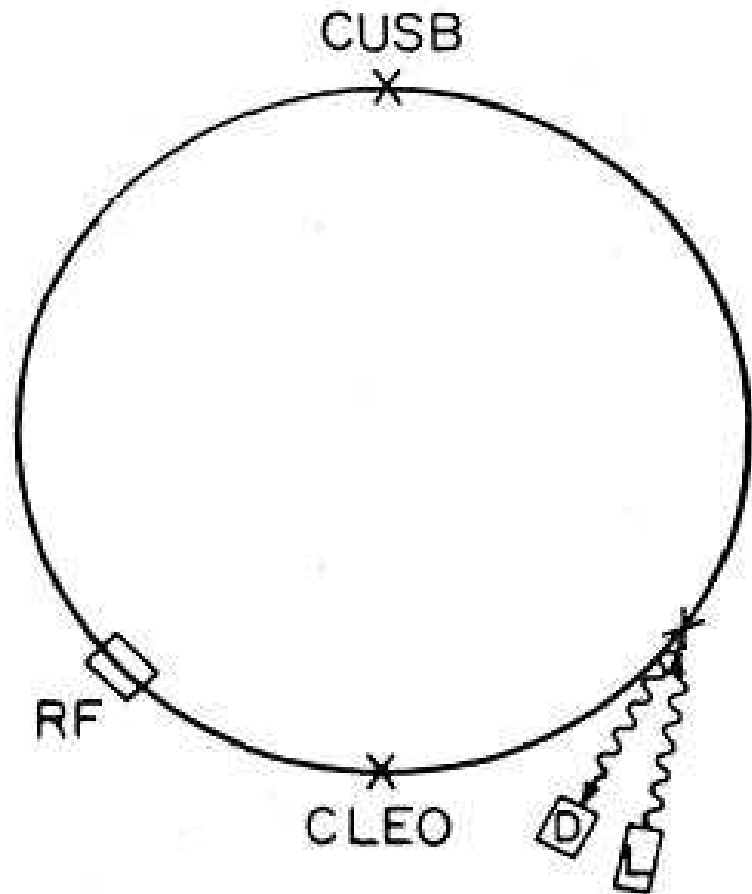
$$U_\gamma = \oint \frac{P_\gamma}{c} ds$$
$$\simeq \frac{C_\gamma U^4}{2\pi} \oint \frac{ds}{\rho^2},$$

$$C_\gamma = \frac{4\pi}{3} \frac{r_e}{(mc^2)^3} = 8.85 \times 10^{-5} \frac{\text{m}}{(\text{GeV})^3}$$

If all dipoles are identical, then $U_\gamma = \frac{C_\gamma U^4}{\rho}$.



CESR circa 1982



$e^+ e^-$ collider.

$L = 768$ m.

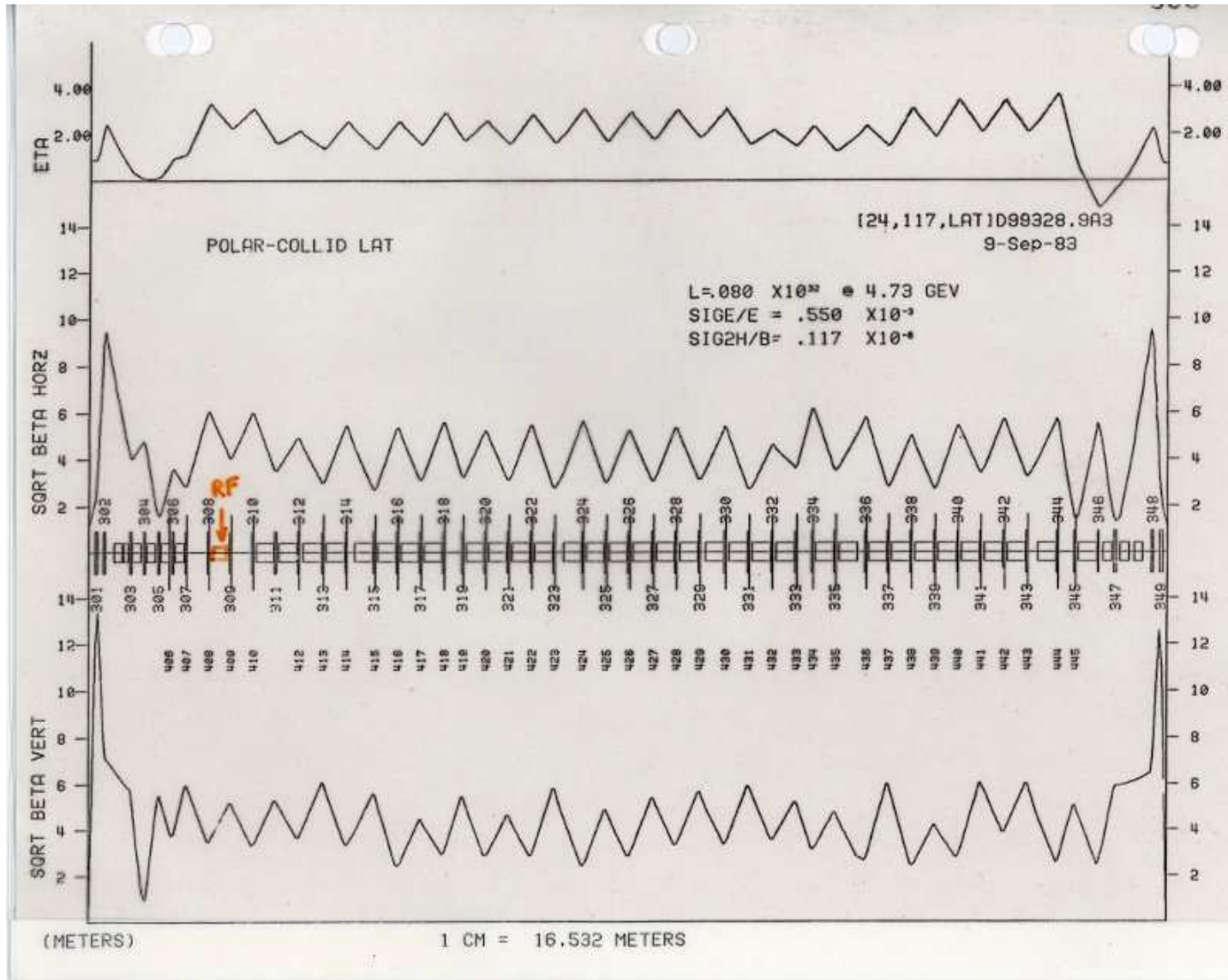
$U \sim 4 \rightarrow 5.6$ GeV.

$f_s = 390$ kHz.

- positrons — clockwise.
- electrons — counterclockwise.



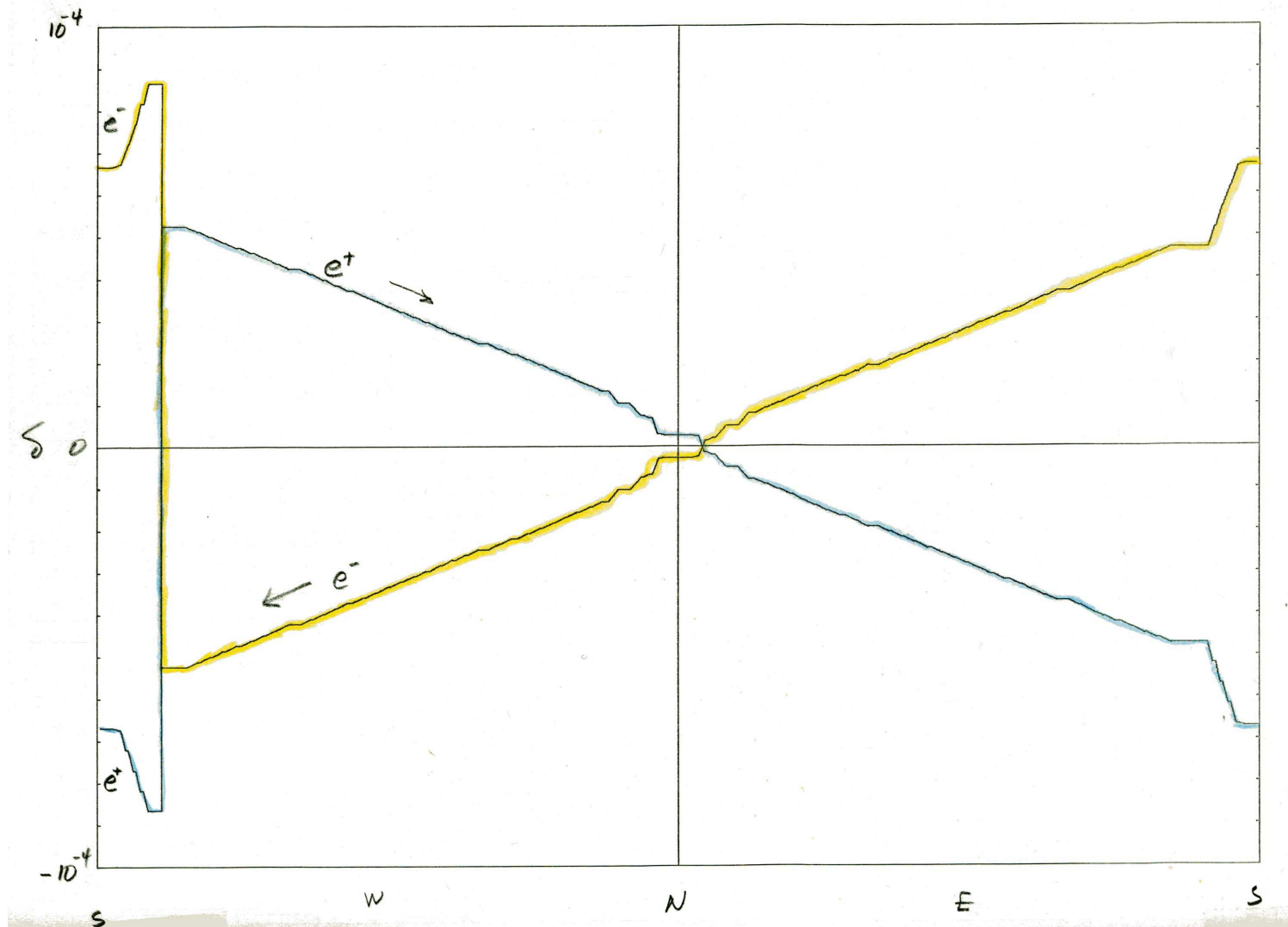
CESR optics (half ring)



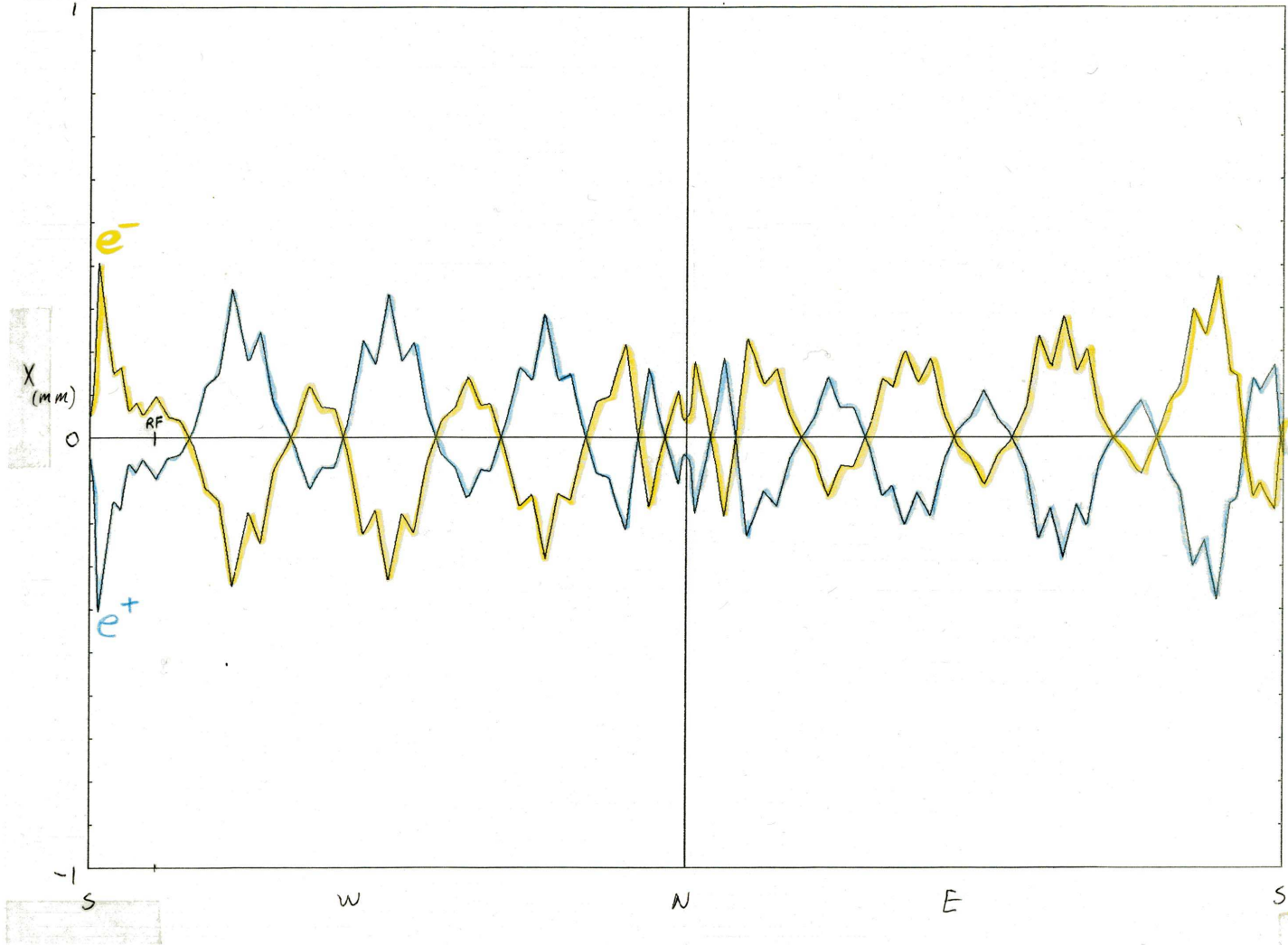
USPAS: Lectures on Radiation & damping
Waldo MacKay January, 2013



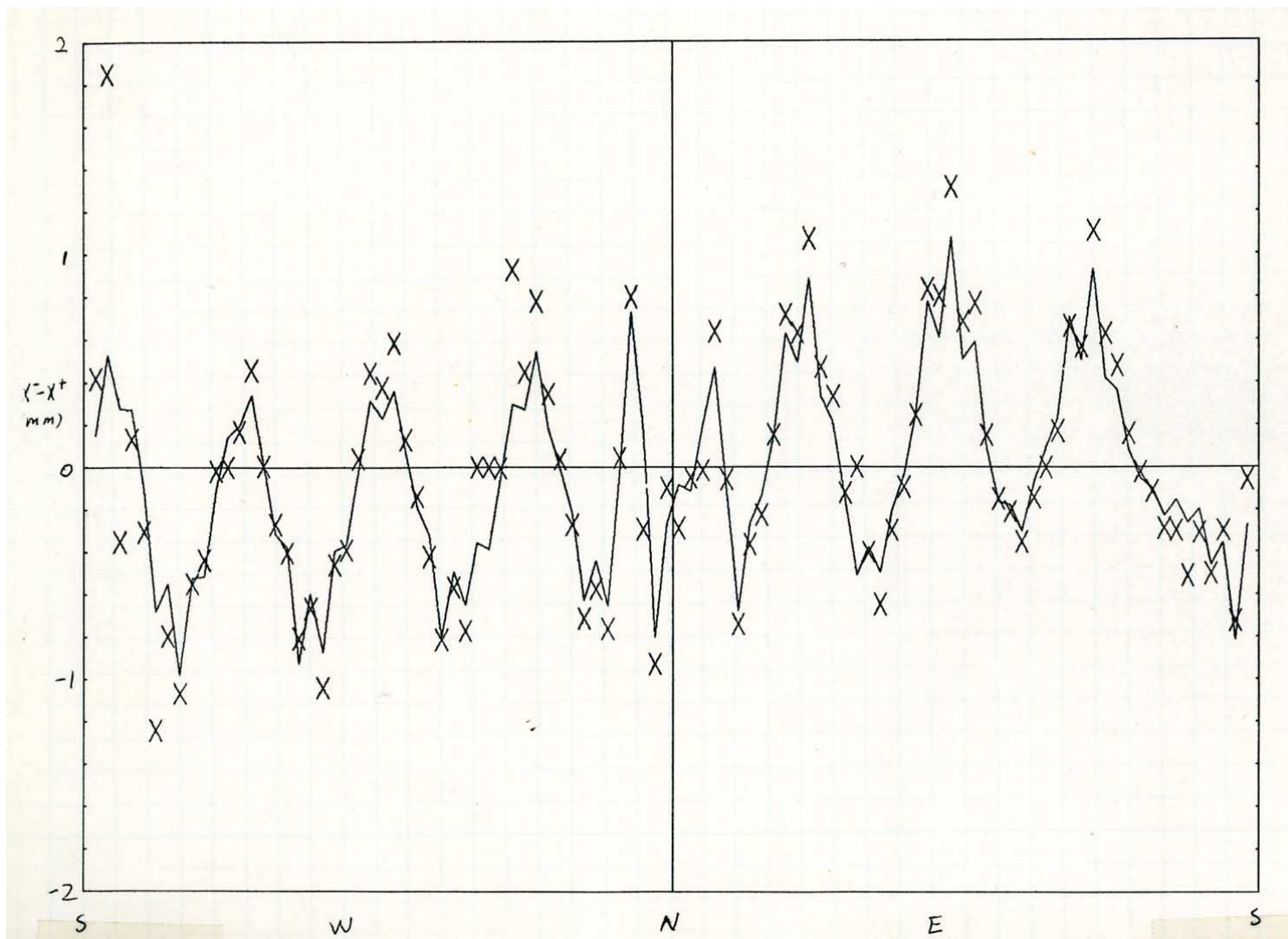
Calculated energy loss



Calculated orbits



Calculated and measured orbit difference



Radiation damping of Energy Oscillations

- energy deviation of particle in question: $u = U - U_s$
- Difference equations of motion:

$$\frac{du}{dt} \simeq \frac{\omega_s}{2\pi} qV \cos \phi_s \varphi - \frac{1}{\tau_s} [U_\gamma(U_s + u) - U_\gamma(U_s)],$$
$$\frac{d\varphi}{dt} \simeq -\frac{\omega_{\text{rf}} \eta_{\text{tr}}}{U_s} u,$$

- Expand energy loss about the synchronous energy U_s :

$$U_\gamma(U) = U_\gamma(U_s) + \left(\frac{dU_\gamma}{dU} \right)_s u + \dots$$

- Combining these yields damped oscillator equation:

$$\frac{d^2 u}{dt^2} + \frac{\omega_s}{2\pi} \left(\frac{dU_\gamma}{dU} \right)_s \frac{du}{dt} + \Omega_s^2 u = 0, \quad \text{with} \quad \Omega_s = \omega_s \sqrt{\frac{qV h \eta_{\text{tr}} \cos \phi_s}{2\pi U_s}}$$

- Of course $\eta_{\text{tr}} = -\alpha_p$ for most electron rings with $\gamma \gg \gamma_{\text{tr}}$.



Damped solution

$$u(t) = u_0 e^{-t/\tau_u} \sin(\Omega'_s t + \psi_0),$$

with damping rate and modified frequency:

$$\frac{1}{\tau_u} = \frac{1}{2\tau_s} \left(\frac{dU_\gamma}{dU} \right)_s \quad \text{and} \quad \Omega'_s = \sqrt{\Omega_s^2 - \frac{1}{\tau_u^2}},$$

The derivative, $(dU_\gamma/dU)_s$, can be calculated from the formula:

$$U_\gamma = \oint P_\gamma \frac{dt}{ds} ds = \frac{1}{c} \oint P_\gamma \left(1 + \frac{\eta}{\rho} \frac{u}{U_s} \right) ds,$$

since the particle has velocity: $c = \frac{d\sigma}{dt} = \left(1 + \frac{\eta}{\rho} \frac{u}{U_s} \right) \frac{ds}{dt}$. (see CM:§5.6)



- Differentiating with respect to energy gives:

$$\left(\frac{dU_\gamma}{dU}\right)_s = \frac{1}{c} \oint \left[\left(\frac{dP_\gamma}{dU}\right)_s + P_\gamma \frac{\eta}{\rho U_s} + \left(\frac{dP_\gamma}{dU}\right)_s \frac{\eta}{\rho U_s} \right] ds.$$

- The 3rd term is negligible: $\Omega_s \gg \frac{1}{\tau_u}$.

e.g. CESR at 5 GeV: $\Omega_s/2\pi \simeq 20$ khz, whereas damping time ~ 10 ms.

- On average: $\frac{dx}{dU} = \frac{\eta}{U_s}$, and $P_\gamma \propto U^2 B^2$, so

$$\left(\frac{dP_\gamma}{dU}\right)_s = 2\frac{P_\gamma}{U_s} + 2\frac{P_\gamma}{\rho} \frac{\rho}{B_0} \left(\frac{\partial B_y}{\partial x}\right)_s \frac{\eta}{U_s}.$$

- Integrating: $\left(\frac{dU_\gamma}{dU}\right)_s = \frac{U_\gamma}{U_s} (2 + \mathcal{D})$, where

$$\mathcal{D} = \frac{1}{cU_\gamma} \oint P_\gamma \eta \frac{1-2n}{\rho} ds, \quad \text{with } n = -\frac{\rho}{B_0} \left(\frac{\partial B_y}{\partial x}\right)_0$$



- Damping time for synchrotron oscillations:

$$\tau_u = \frac{2}{2 + \mathcal{D}} \frac{U_s}{U_\gamma} \tau_s.$$

- For separated function lattice ($n = 0$ in dipoles), we should note that

$$\mathcal{D} \simeq \frac{1}{cU_\gamma} \oint P_\gamma \frac{\eta}{\rho} ds.$$



Damping of vertical oscillations

- trajectory: $y = \sqrt{\mathcal{W}\beta} \cos \psi$
- Courant-Snyder invariant: $\mathcal{W} = \beta y'^2 + 2\alpha y y' + \gamma y^2$
- For simplicity (following Matt Sands), let's set $\alpha = 0$, $\Rightarrow \gamma = 1/\beta$.
(The gory details with nonconstant α are worked out in the book and end up giving the same result.)
- So we have for a given trajectory:

$$y = A\sqrt{\beta} \cos \psi, \quad \text{and} \quad y' = -\frac{A}{\sqrt{\beta}} \sin \psi$$

$$\mathcal{W} = \beta y'^2 + \frac{y^2}{\beta} = \beta \frac{A^2}{\beta} \sin^2 \psi + \frac{1}{\beta} A^2 \beta \cos^2 \psi = A^2$$



- For damping, we ignore quantum fluctuations and take the momentum loss in the direction of motion:
- When a photon is radiated, the electron loses Δu of energy, but: $\Delta y = 0$ (always), and $\Delta y' = 0$ (ignoring quant. fluct.)
- before the photon is radiated: $p_{y0} = p_0 y'_0$.
- after the photon is radiated: $p_{y1} = p_0 y'_0 \left(1 - \frac{\Delta u}{U_s}\right)$
- integrating the loss around the ring gives: $p_y \simeq p_0 y' \left(1 - \frac{U_\gamma}{U_s}\right)$
- longitudinal component after one turn is $p_z \simeq p_0 \left(1 - \frac{U_\gamma}{U_s}\right)$
- add rf kick to recover energy: $p_z \simeq p_0 \left(1 - \frac{U_\gamma}{U_s}\right) \left(1 + \frac{U_\gamma}{U_s}\right)$
- after complete revolution: $y'_{\text{rf}} \simeq y' \left(1 - \frac{U_\gamma}{U_s}\right)$.



- per turn we have: $\Delta y' = -y' \frac{U_\gamma}{U_s}$, and $\Delta y = 0$.
- How does our invariant change?

$$(A + \Delta A)^2 - A^2 = 2A \Delta A + (\Delta A)^2 \simeq 2A \Delta A,$$

if we assume the $(\Delta A)^2$ term is **small** compared to $2A \Delta A$.

$$2A \Delta A \simeq \beta(y'_0 + \Delta y')^2 + \frac{(y_0 + 0)^2}{\beta} - \left(\beta y_0'^2 + \frac{y_0^2}{\beta} \right)$$

$$\simeq 2\beta y'_0 \Delta y' + \beta (\Delta y')^2 \simeq 2\beta y'_0 \Delta y'$$

$$A \Delta A \simeq -\beta y_0'^2 \frac{U_\gamma}{U_s}$$

$$\simeq -\beta \left\langle \left(\frac{A}{\sqrt{\beta}} \sin \psi \right)^2 \right\rangle_\psi \frac{U_\gamma}{U_s} \quad (\text{average over all betatron-phases})$$

$$= \frac{1}{2} A^2 \frac{U_\gamma}{U_s}$$



$$\frac{dA}{dt} \simeq \frac{\Delta A}{\tau_s} \simeq -\frac{1}{2} \frac{U_\gamma}{U_s \tau_s} A$$

$$A = A_0 e^{-t/\tau_y} \quad \text{with}$$

$$\tau_y = 2 \frac{U_s}{U_\gamma} \tau_s = 2 \frac{U_s}{\langle P_\gamma \rangle},$$

where $\langle P_\gamma \rangle$ radiated power averaged over one turn.



Damping or horizontal oscillations

- trajectory equations:

$$x = x_\beta + x_p, = x_\beta + \eta \frac{u}{U_s},$$

$$x_\beta = A\sqrt{\beta} \cos \psi,$$

$$x' = x'_\beta + x'_p = x'_\beta + \eta' \frac{u}{U_s},$$

$$x'_\beta = \frac{A}{\sqrt{\beta}} \sin \psi, \quad (\alpha = 0 \text{ approx.})$$

- When electron radiates, $dx = 0$ and $dx' = 0$ just like vertical, but

$$dx_\beta = -\eta \frac{-du}{U_s}, \quad \text{and} \quad dx'_\beta = -\eta' \frac{-du}{U_s}$$



- Courant-Snyder invariant: $\mathcal{W} = A^2 = \beta x'_\beta{}^2 + \frac{x_\beta^2}{\beta}$

$$2A d(\Delta A) + [d(\Delta A)]^2 = \beta(x'_\beta + dx'_\beta)^2 + \frac{1}{\beta}(x_\beta + dx_\beta)^2 - \left(\beta x'_\beta{}^2 + \frac{x_\beta^2}{\beta} \right)$$

$$\simeq 2\beta x'_\beta dx'_\beta + 2x_\beta dx_\beta + \mathcal{O}(d^2)$$

$$A d(\Delta A) = \beta \eta' \frac{dU}{U_s} x'_\beta + \frac{1}{\beta} x_\beta \eta \frac{dU}{U_s}$$

$$= \left(\beta \eta' x'_\beta + \frac{\eta x_\beta}{\beta} \right) \frac{dU}{U_s}$$



Recall: $P_\gamma \propto U^2 B^2$, so

$$\begin{aligned}
 \frac{dU_\gamma}{ds} &= \frac{1}{c} \frac{dU_\gamma}{dt} \left(1 + \frac{x}{\rho}\right) = \frac{P_\gamma}{c} \left(1 + \frac{2}{B} \frac{\partial B_y}{\partial x} x\right) \left(1 + \frac{x}{\rho}\right) \\
 &= \frac{P_\gamma}{c} \left(1 - \frac{2n}{\rho} x\right) \left(1 + \frac{x}{\rho}\right) \\
 &= \frac{P_\gamma}{c} \left[1 - \frac{2n}{\rho} (x_\beta + x_p)\right] \left(1 + \frac{x_\beta + x_p}{\rho}\right) \\
 &= \frac{P_\gamma}{c} \left[1 + \frac{1-2n}{\rho} x_\beta - \frac{4n}{\rho^2} x_\beta x_p - \frac{2n}{\rho^2} x_\beta^2 + \left(1 + \frac{x_p}{\rho} - \frac{2n}{\rho^2} x_p^2\right)\right]
 \end{aligned}$$

Average over all betatron and synchrotron phases:

$$\langle A d(\Delta A) \rangle_\psi \rightarrow \frac{P_\gamma}{cU_s} \left\langle \left(\beta\eta' x'_\beta + \frac{\eta x_\beta}{\beta} \right) x_\beta \left[\frac{1-2n}{\rho} - \frac{4n}{\rho^2} \frac{u}{U_s} \right] \right\rangle_\psi ds$$



Since

$$\langle x'_\beta x_\beta \rangle = 0, \quad \text{and} \quad \langle x_\beta^2 \rangle = \frac{A^2 \beta}{2},$$

we find

$$\langle d(\Delta A) \rangle_\psi = \frac{A}{2cU_s} P_\gamma \eta \frac{1 - 2n}{\rho} ds,$$

and integrating around the ring, the contribution from radiation is

$$(\Delta A)_{\text{rad}} = \frac{A}{2cU_s} \oint P_\gamma \eta \frac{1 - 2n}{\rho} ds = \frac{A}{2} \frac{U_\gamma}{U_s} \mathcal{D}$$



- Now for the rf contribution to ΔA .
Just in front of the cavity:

$$x_1 = x_{\beta 1} + \eta \frac{u_1}{U_s}, \quad \text{and} \quad x'_1 = x'_{\beta 1} + \eta' \frac{u_1}{U_s}$$

Right after the cavity we have

$$x_2 = x_{\beta 2} + \eta \frac{u_2}{U_s} = x_1, \quad \text{and} \quad x'_2 = x'_{\beta 2} + \eta' \frac{u_2}{U_s} = \left(1 - \frac{U_\gamma}{U_s}\right) x'_1.$$

So with $u_2 = u_1 + U_\gamma$ from the rf kick

$$(\Delta x_\beta)_{\text{rf}} = x_{\beta 2} - x_{\beta 1} = -\eta \frac{U_\gamma}{U_s},$$

$$x'_2 - x'_1 = \Delta x_\beta + \eta' \frac{U_\gamma}{U_s} = -\frac{U_\gamma}{U_s} x'_1, \quad \text{or on rearranging}$$

$$(\Delta x'_\beta)_{\text{rf}} = x'_{\beta 2} - x'_{\beta 1} = -x_{\beta 1} \frac{U_\gamma}{U_s} - \eta' \frac{U_\gamma}{U_s} - \eta' \frac{U_\gamma u_1}{U_s^2}$$



For the variation of the amplitude, we have the first order terms:

$$2A \Delta A = 2 \left(\beta x'_\beta \Delta x'_\beta + \frac{x_\beta \Delta x_\beta}{\beta} \right) + \mathcal{O}(\Delta^2),$$

but averaging over all betatron phases:

$$\langle x_\beta \Delta x_\beta \rangle \propto \langle x_\beta \rangle \rightarrow 0, \quad \text{and} \quad \langle x'_\beta \Delta x'_\beta \rangle \rightarrow -\frac{U_\gamma}{U_s} \langle x'^2_\beta \rangle = \frac{A^2}{2\beta} \frac{U_\gamma}{U_s}.$$

$$A(\Delta A)_{\text{rf}} = -\frac{U_\gamma}{U_s} \beta \langle x'^2_\beta \rangle = -\frac{A^2}{2} \frac{U_\gamma}{U_s}$$

- Adding the radiation and rf parts together yields:

$$\Delta A = (\Delta A)_{\text{rf}} + (\Delta A)_{\text{rad}} = -\frac{A}{2} \frac{U_\gamma}{U_s} + \frac{A}{2} \frac{U_\gamma}{U_s} \mathcal{D}$$



Dividing by the revolution period, we get the differential equation:

$$\frac{dA}{dt} = \frac{\Delta A}{\tau_s} = -\frac{U_\gamma}{2U_s\tau_s}(1 - \mathcal{D}) A$$

which has the solution

$$A = A_0 e^{-t/\tau_x},$$

with damping time

$$\tau_x = \frac{2}{1 - \mathcal{D}} \frac{U_s}{U_\gamma} \tau_s.$$



To summarize the damping rates, define

$$\frac{1}{\tau_0} = \frac{U_\gamma}{2U_s\tau_s}.$$

$$\frac{1}{\tau_x} = \frac{1}{\tau_0} (1 - \mathcal{D}),$$

$$\frac{1}{\tau_y} = \frac{1}{\tau_0},$$

$$\frac{1}{\tau_u} = \frac{1}{\tau_0} (2 + \mathcal{D}).$$

The sum of the three rates is

$$\frac{1}{\tau_x} + \frac{1}{\tau_y} + \frac{1}{\tau_u} = \frac{4}{\tau_0}.$$

Partition numbers: $J_x = 1 - \mathcal{D}$, $J_y = 1$ and $J_u = 2 + \mathcal{D}$

$$\sum_i J_i = 4$$



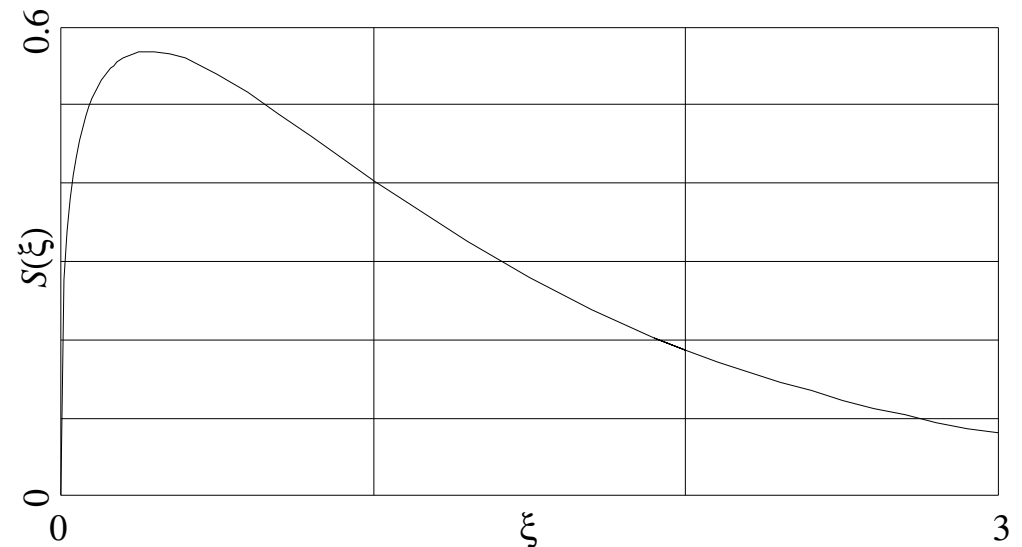
Some Properties of Synchrotron Radiation

- Power spectrum:

$$dP_\gamma = \frac{P_\gamma}{\omega_c} S\left(\frac{\omega}{\omega_c}\right) d\omega$$

$$S(\xi) = \frac{9\sqrt{3}}{8\pi} \xi \int_\xi^\infty K_{\frac{5}{3}}(\xi') d\xi'$$

$$\int_0^\infty S(\xi) d\xi = 1$$



- Critical energy: half the power is radiated by photons less than the critical energy, and the other half, above.

$$u_c = \hbar\omega_c = \frac{3\hbar c}{2\rho} \gamma^3$$



- Number of photons per second:

$$N_\gamma = \int_0^{U_{\max}} n_\gamma(u_\gamma) du_\gamma = \frac{5}{2\sqrt{3}} \frac{\alpha_f c}{\rho} \gamma$$

here: $\alpha_f = 1/137$

- Number of photons per radian:

$$N_r = \frac{5\alpha_f}{2\sqrt{3}} \gamma$$

- Average photon energy and 2nd moment:

$$\langle u_\gamma \rangle = \frac{1}{N_\gamma} \int_0^{U_{\max}} u n_\gamma(u) du = \frac{8}{15\sqrt{3}} u_c \simeq 0.32 u_c$$

$$\langle u_\gamma^2 \rangle = \frac{1}{N_\gamma} \int_0^{U_{\max}} u^2 n_\gamma(u) du = \frac{11}{27\sqrt{3}} u_c^2 \simeq 0.41 u_c^2$$

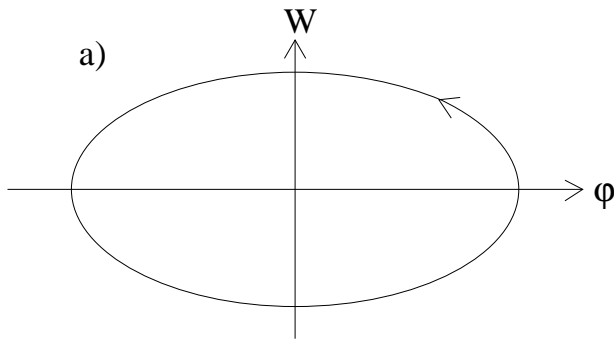


Table I. Estimate of radiation for several rings.

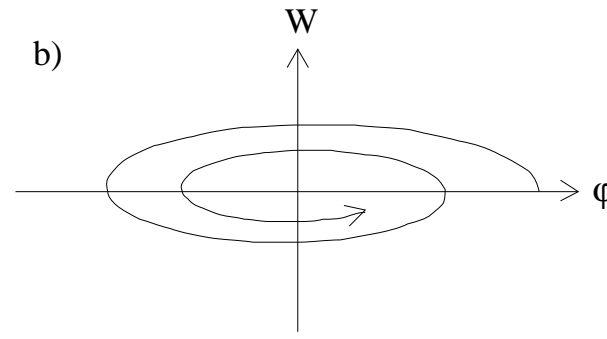
Machine	U_s [GeV]	$\langle u_\gamma \rangle$ [eV]	N_γ [s ⁻¹]	N_r
Adone	1.5	4.6×10^2	1.8×10^9	31
SPEAR	4.5	4.9×10^3	2.2×10^9	93
CESR	5.0	5.0×10^2	4.0×10^8	160
PETRA	19	2.5×10^4	6.1×10^8	390
LEP I	50	1.4×10^4	5.0×10^7	1030
LEP II	100	1.1×10^5	1.0×10^8	2060



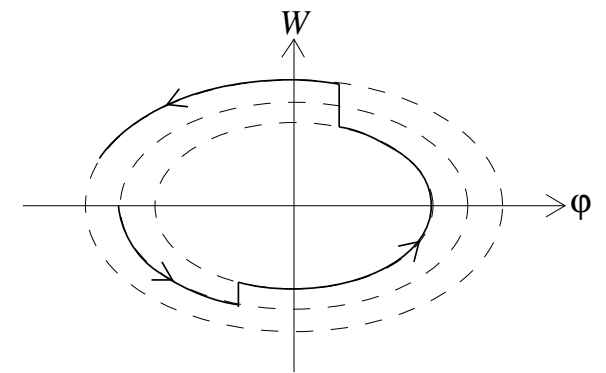
Energy spread and bunch length



No radiation



Continuous damping : Quantum fluctuation effect



- Equilibrium: damping rate = growth rate from quantum excitations.

- Recall the longitudinal phase space variables: W and φ

$$W = -\frac{u}{\omega_{\text{rf}}} = -\frac{U_s}{\omega_{\text{rf}}^2 \alpha_p} \frac{d\varphi}{dt}, \quad (\eta_{\text{tr}} \rightarrow -\alpha_p \text{ for ultrarelativistic electrons.})$$



- Longitudinal phase space variables:

$$u = -\omega_{\text{rf}}W = -A \cos[\Omega_s(t - t_0)],$$

$$\varphi = \frac{\omega_{\text{rf}}\alpha_p}{U_s\Omega_s} A \sin[\Omega_s(t - t_0)].$$

For simplicity, rescale the rf phase to

$$\xi = \frac{U_s\Omega_s}{\omega_{\text{rf}}\alpha_p} \varphi.$$

Then we have the Courant-Snyder type invariant:

$$\xi^2 + u^2 = A^2.$$

- When a photon of energy u_γ is emitted, this invariant changes by

$$\Delta(A^2) = [(u - u_\gamma)^2 + \xi^2] - [u^2 + \xi^2] = -2uu_\gamma + u_\gamma^2.$$



- Summing over all emissions around the ring:

$$\Delta(A^2) \simeq -2u \oint N_\gamma \langle u_\gamma \rangle \frac{ds}{c} + \oint N_\gamma \langle u_\gamma^2 \rangle \frac{ds}{c}.$$

- averaging over synchrotron phase: $\langle u_\gamma \rangle = 0$.

$$\left(\frac{d(A^2)}{dt} \right)_{\text{QF}} \simeq \frac{\Delta(A^2)}{\tau_s} = \frac{1}{c\tau_s} \oint N_\gamma \langle u_\gamma^2 \rangle ds$$

- Damping: $A = A_0 e^{-t/\tau_u} \Rightarrow \left(\frac{d(A^2)}{dt} \right)_{\text{damping}} = -\frac{2}{\tau_u} A^2$

- Equilibrium: $0 = -\frac{2}{\tau_u} \langle A^2 \rangle + \frac{1}{c\tau_s} \oint N_\gamma \langle u_\gamma^2 \rangle ds$

$$\sigma_u^2 = \frac{\langle A^2 \rangle}{2} = \frac{\tau_u}{4c\tau_s} \oint N_\gamma \langle u_\gamma^2 \rangle ds.$$



- Energy spread:

$$\sigma_u = \sqrt{\frac{C_q}{J_u \rho}} \gamma^2 mc^2$$

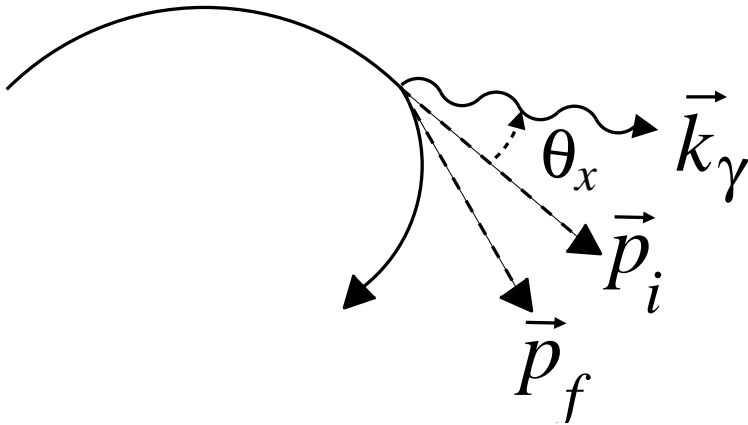
with $C_q = 3.8 \times 10^{-8}$ m and $J_u \sim 2 + \mathcal{D}$.

Table II. Energy spread of several rings.

Machine	U_s [GeV]	σ_u [MeV]	σ_u/U_s
Adone	1.5	0.86	5.7×10^{-4}
SPEAR	4.5	4.9	1.1×10^{-3}
CESR	5.0	3.0	5.8×10^{-4}
PETRA	19	27.3	1.2×10^{-3}
LEP I	50	55	1.1×10^{-3}
LEP II	100	220	2.2×10^{-3}



Transverse excitations and emittances



$$k_\gamma = \frac{u_\gamma}{c}$$

$$\vec{p}_f = \vec{p}_i - \vec{k}_\gamma$$

$$\langle \theta_x^2 \rangle \sim \frac{1}{\gamma^2}$$

- Projection of recoil angle:

$$\Delta p_x \simeq -\theta_x u_\gamma$$

$$\Delta x' \simeq \frac{\Delta p_x}{p} \simeq -\theta_x \frac{u_\gamma}{U_s}$$

$$\delta x_\beta = \eta \frac{u_\gamma}{U_s},$$

$$\delta x'_\beta = (\eta' - \theta_x) \frac{u_\gamma}{U_s},$$



- For horizontal the dispersion ($\eta u/U_s$) typically dominates the $\langle \theta_x^2 \rangle$.
 e.g, $\eta \sim 1$ to 2 m with quad separation of 10 m, $\Rightarrow \eta' \sim 0.05$ to 0.1.
 1 GeV electrons have $\gamma \simeq 2 \times 10^3$ and $\theta_{x,\text{rms}} \sim 5 \times 10^{-4} \ll \eta'$

$$\begin{aligned} \Delta(A^2) &= \beta_{\text{H}}(x'_{\beta} + \delta x'_{\beta})^2 + 2\alpha_{\text{H}}(x_{\beta} + \delta x_{\beta})(x'_{\beta} + \delta x'_{\beta}) + \gamma_{\text{H}}(x_{\beta} + \delta x_{\beta})^2 \\ &\quad - [\beta_{\text{H}}x_{\beta}^{\prime 2} + 2\alpha_{\text{H}}x_{\beta}x'_{\beta} + \gamma_{\text{H}}x_{\beta}^2] \\ &= 2[\beta_{\text{H}}x'_{\beta} \delta x'_{\beta} + \alpha_{\text{H}}(x_{\beta}\delta x'_{\beta} + x'_{\beta}\delta x_{\beta}) + \gamma_{\text{H}}x_{\beta}\delta x_{\beta}] && \text{(Line 1)} \\ &\quad + \beta_{\text{H}}(\delta x'_{\beta})^2 + 2\alpha_{\text{H}}\delta x_{\beta}\delta x'_{\beta} + \gamma_{\text{H}}(\delta x_{\beta})^2 && \text{(Line 2)} \end{aligned}$$

- Line 1 is linear in u_{γ} and θ_x .
 - The linear θ_x terms average to zero.
 - The linear u_{γ} term just averages to give the previous damping contribution.
- Line 2 is quadratic in u_{γ} and θ_x and will give positive averages.



$$\Delta(A^2) = \left(\frac{u_\gamma}{U_s} \right)^2 [\beta_H \eta'^2 + 2\alpha_H \eta \eta' + \gamma_H \eta^2] + \frac{1}{\gamma^2} \frac{\langle u_\gamma^2 \rangle}{U_s^2} \beta_H$$

- It's usual to define:

$$\mathcal{H}(s) = \beta_H(s) (\eta'(s))^2 + 2\alpha_H(s) \eta(s) \eta'(s) + \gamma_H(s) (\eta(s))^2.$$

- Some people mistakenly refer to \mathcal{H} as “the dispersion invariant”.
 - $\mathcal{H}(s)$ remains constant in straight sections, but
 - $\mathcal{H}(s)$ varies in bends.
- I call it the “curly H-function”.
- Excitation rate:

$$\left(\frac{d(A^2)}{dt} \right)_{\text{QF}} \simeq \frac{1}{c\tau_s} \oint \frac{\langle u_\gamma^2 \rangle}{U_s} \mathcal{H}(s) N_\gamma ds.$$



For equilibrium, sum of the damping and excitation rates must be zero:

$$-\left(\frac{d(A^2)}{dt}\right)_{\text{damping}} = \frac{2}{\tau_x} \langle A^2 \rangle \simeq \frac{1}{c\tau_s} \oint \frac{\langle u_\gamma^2 \rangle}{U_s^2} \mathcal{H}(s) N_\gamma ds.$$

or

$$\langle A^2 \rangle = \frac{\tau_x}{2\tau_s c} \oint \frac{\langle u_\gamma^2 \rangle}{U_s^2} \mathcal{H}(s) N_\gamma ds.$$

$$\sigma_{x_\beta}^2 = \frac{1}{2} \beta \langle A^2 \rangle = \beta_H \epsilon_{\text{rms}},$$

with

$$\epsilon_{\text{rms}} = \frac{\tau_x}{4LU_s^2} \oint \langle u_\gamma^2 \rangle \mathcal{H}(s) N_\gamma ds.$$

- In high energy e^\pm rings, it is more usual to quote rms emittances which are unnormalized.



- For flat rings, vertical dispersion is zero, therefore the vertical emittance is typically much smaller than the horizontal.

$$\langle \Delta(A^2) \rangle \simeq \frac{1}{\gamma^2} \frac{u_\gamma^2}{U_s^2} \beta_V.$$

- In most rings, there is a little xy -coupling, so the vertical emittance is frequently dominated by the coupling.
- ϵ_V/ϵ_H may be only a few percent or less in a well aligned light source, but in e^+e^- colliders, the ratio is frequently 10—20%.
- For low emittance lattices: minimize \mathcal{H} .
 - FODO lattices not optimized.
 - Chasman-Green style lattices with double bend or tripple-bend achromats are much better fro low emittance.



Emittance as temperature (see §12.1)

- In the rest system of the design particle (average CM of beam):

$$\beta_{\perp}^* = \frac{p_{\perp}^* c}{U^*} = \frac{p_{\perp} c}{U/\gamma_0} = \gamma_0 \beta_{\perp}.$$

Note: $U^* \simeq mc^2$ is practically at rest ($\beta^* \ll 1$).

- Recall from thermodynamics: $\frac{1}{2}k_B T$ for each degree of freedom.
 - $\langle p_{\perp}^2 \rangle = \langle p_x^2 + p_y^2 \rangle = p^2 (\sigma_{x'}^2 + \sigma_{y'}^2)$
- The transverse temperature T_{\perp} of the beam may be defined as

$$k_B T_{\perp} = \left\langle \frac{1}{2} m v_{\perp}^{*2} \right\rangle = \left\langle \frac{1}{2} m c^2 \beta_{\perp}^{*2} \right\rangle = \left\langle \frac{1}{2} m c^2 \gamma_0^2 \beta_{\perp}^2 \right\rangle$$



$$\beta_{\perp} \simeq \beta_0 \frac{p_{\perp}}{p}$$

$$k_B T_{\perp} = \left\langle \frac{1}{2} m c^2 (\beta_0 \gamma_0)^2 \frac{p_{\perp}^2}{p^2} \right\rangle$$

$$\langle p_{\perp}^2 \rangle = p^2 \left(\frac{\epsilon_{H,\text{rms}}}{\beta_H} + \frac{\epsilon_{V,\text{rms}}}{\beta_V} \right)$$

$$k_B T_{\perp} = \frac{1}{2} m c^2 (\beta_0 \gamma_0)^2 \left(\frac{\epsilon_{H,\text{rms}}}{\beta_H} + \frac{\epsilon_{V,\text{rms}}}{\beta_V} \right), \quad \text{unnormalized emittance}$$

$$= \frac{1}{2} m c^2 (\beta_0 \gamma_0) \left(\frac{\epsilon_{H,\text{rms}}^*}{\beta_H} + \frac{\epsilon_{V,\text{rms}}^*}{\beta_V} \right), \quad \text{normalized emittance}$$

- So $\frac{k_B T_{\perp}}{\beta_0 \gamma_0}$ is an invariant.



- Longitudinal temperature: $\frac{1}{2}k_B T_{\perp} = \left\langle \frac{1}{2}mc^2 \beta_{\parallel}^{*2} \right\rangle$

- Boosting β_{\parallel}^* from the lab system: $\beta_{\parallel}^* = \frac{\beta_{\parallel} - \beta_0}{1 - \beta_{\parallel}\beta_0}$

- Define

$$\Delta\beta = \beta_{\parallel} - \beta_0$$

$$\Delta\gamma = \gamma - \gamma_0$$

- Recall

$$\gamma = (1 - \beta^2)^{-1/2}$$

$$\Delta\gamma \simeq \beta(1 - \beta^2)^{-3/2} \Delta\beta \simeq \beta_0 \gamma_0^3 \Delta\beta$$

$$\beta_{\parallel}^* = \frac{\Delta\beta}{1 - (\beta_0 + \Delta\beta)\beta_0} \simeq \frac{\Delta\beta}{1 - (\beta_0 + \Delta\beta)\beta_0} \simeq \frac{\Delta\beta}{\gamma_0^{-2} - \gamma_0^{-3} \Delta\gamma}$$

$$\simeq \gamma_0^2 \Delta\beta, \quad \text{assuming } \frac{\Delta\gamma}{\gamma_0} \simeq \frac{\Delta p}{p} \ll 1.$$

$$\simeq \frac{\Delta\gamma}{\beta_0 \gamma_0},$$



$$\begin{aligned} \frac{1}{2} k_B T_{\perp} &\simeq \frac{1}{2} mc^2 \left\langle \left(\frac{\Delta\gamma}{\beta_0 \gamma_0} \right)^2 \right\rangle \simeq \frac{1}{2} mc^2 \beta_0^2 \left\langle \left(\frac{\Delta p}{p} \right)^2 \right\rangle, \\ &\simeq \frac{1}{2} mc^2 \beta_0^2 \left(\frac{\sigma_p}{p} \right)^2 \end{aligned}$$

since

$$\frac{dU}{U} = \beta^2 \frac{dp}{p}.$$

One is tempted to write for the overall temperature:

$$\frac{3}{2} k_B T = \frac{mc^2}{2} (\beta_0 \gamma_0)^2 \left[\frac{\epsilon_{H,\text{rms}}}{\beta_H} + \frac{\epsilon_{V,\text{rms}}}{\beta_V} + \frac{1}{\gamma_0^2} \left(\frac{\sigma_p}{p} \right)^2 \right].$$

