

# Basics of Polarized Beam Acceleration

## Protons:

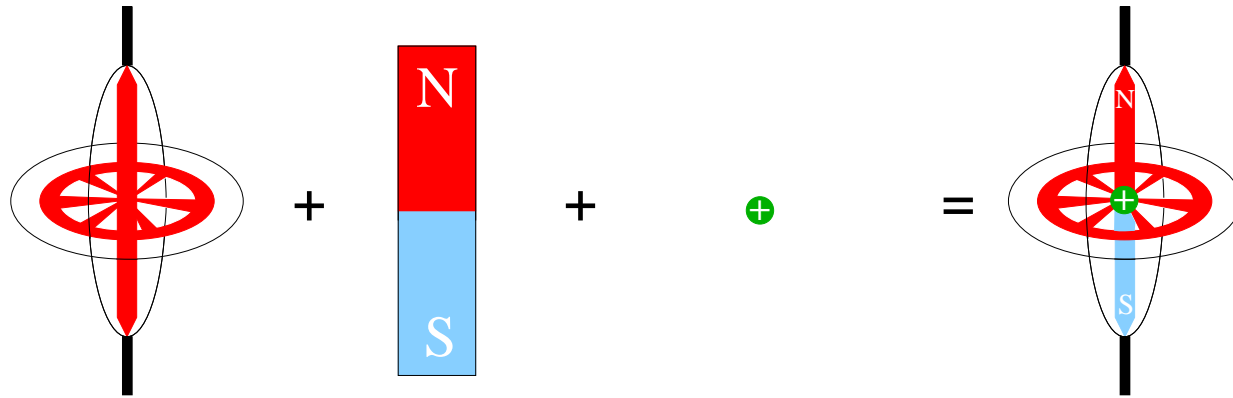
- Simple model of the proton.
  - Spin dynamics.
  - Depolarizing resonances.
  - Siberian snakes.
- The real machines: RHIC and injectors.

## Electrons/Positrons:

- Longitudinal beam dynamics.
  - Synchrotron oscillations and tune.
  - Electrons: Synchrotron radiation
- Radiative polarization.
- Quantum fluctuations  $\Rightarrow$  Spin Diffusion
- Polarization in some real  $e^\pm$  machines.
- Measurements with polarized  $e^\pm$  beams.



# Simple Model of Proton



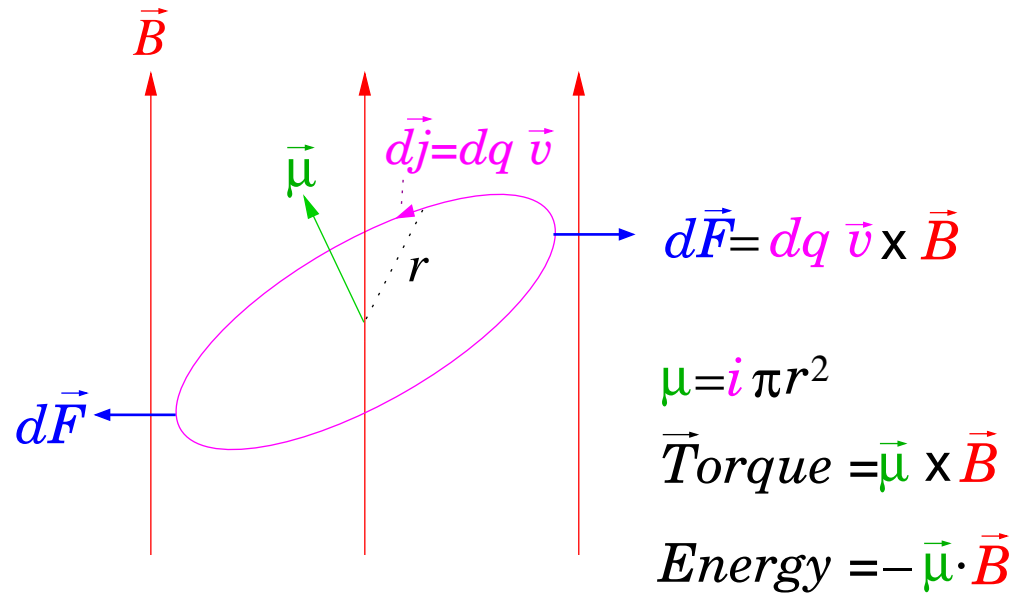
Gyroscope + Bar magnet + Charge = "proton"

Spin  
Magnetic  
Dipole  
Moment

**Polarization: Average spin of the ensemble of protons.**

$$\vec{P} = \frac{1}{N} \sum_{j=1}^N \frac{\vec{S}_j}{|S|}$$

# Torque on Classical Magnetic Moment



~ Semiclassical model:

- The spin  $\vec{S}$  has a constant magnitude in the rest frame.
- The magnetic moment  $\vec{\mu} \propto \vec{S}$ .
  - $\vec{\mu}$  has a constant magnitude in the rest frame.  
(Sort of like a loop of infinite inductance.)



Mass and Charge:

$$m = \int \rho_m d^3 r, \quad \text{and} \quad q = \int \rho_e d^3 r.$$

Magnetic Moment and Spin (intrinsic angular momentum):

$$\vec{\mu} = \frac{1}{2} \int \vec{r} \times \vec{j}(\vec{r}) d^3 r = \frac{1}{2} \int \vec{r} \times \rho_e(\vec{r}) \vec{v}(\vec{r}) d^3 r.$$
$$\vec{S} = \int \vec{r} \times \rho_m(\vec{r}) \vec{v}(\vec{r}) d^3 r.$$

$$\frac{\mu}{S} \sim \frac{1}{2} \frac{q}{m}.$$

(see e. g., Panofsky and Phillips or other E&M text.)



# Relativistic Angular Momentum

Energy-momentum tensor (à la Weinberg)

$$T^{\alpha\beta}(x) = T^{\beta\alpha}(x) = \sum_n \frac{p_n^\alpha p_n^\beta}{E_n} \delta^3(x - x_n(t))$$

For isolated system

$$\frac{\partial}{\partial x^\alpha} T^{\alpha\beta} = 0.$$

Define 4d analogue of  $\vec{r} \times \vec{p}$ :

$$M^{\alpha\beta\gamma} = x^\alpha T^{\beta\gamma} - x^\beta T^{\alpha\gamma}$$

$$J^{\alpha\beta} = \int M^{0\alpha\beta} d^3x = \int x^\alpha T^{\beta 0} - x^\beta T^{\alpha 0} d^3x$$

Spin (intrinsic angular momentum):

$$S_\alpha = \frac{1}{2c} \epsilon_{\alpha\beta\gamma\delta} J^{\beta\gamma} u^\delta, \quad \text{proper velocity: } u^\delta = \frac{dx^\delta}{d\tau}.$$



For a particle at rest with CM at rest at the origin:

$$J^{\diamond\mu\nu} : \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & S_z^\diamond & -S_y^\diamond \\ 0 & -S_z^\diamond & 0 & S_x^\diamond \\ 0 & S_y^\diamond & -S_x^\diamond & 0 \end{pmatrix}, \quad (\vec{J}^\diamond = \vec{S}^\diamond)$$

Boost along  $z$ :

$$J^{\mu\nu} : \begin{pmatrix} 0 & \gamma\beta S_y^\diamond & -\gamma\beta S_x^\diamond & 0 \\ -\gamma\beta S_y^\diamond & 0 & S_z^\diamond & -\gamma S_y^\diamond \\ \gamma\beta S_x^\diamond & -S_z^\diamond & 0 & \gamma S_x^\diamond \\ 0 & \gamma S_y^\diamond & -\gamma S_x^\diamond & 0 \end{pmatrix}, \quad \Rightarrow \quad \vec{J} = \begin{pmatrix} \gamma S_x^\diamond \\ \gamma S_y^\diamond \\ S_z^\diamond \end{pmatrix}$$

$$S^\mu : \begin{pmatrix} \gamma\beta S_z^\diamond \\ S_x^\diamond \\ S_y^\diamond \\ \gamma S_z^\diamond \end{pmatrix}, \quad \Rightarrow \quad \vec{S} = \begin{pmatrix} S_x^\diamond \\ S_y^\diamond \\ \gamma S_z^\diamond \end{pmatrix}, \quad S^0 = \vec{\beta} \cdot \vec{S}$$

$$\vec{J} - \vec{S} = \begin{pmatrix} (\gamma - 1)S_x^\diamond \\ (\gamma - 1)S_y^\diamond \\ (1 - \gamma)S_z^\diamond \end{pmatrix}$$



# Center-of-Mass shift

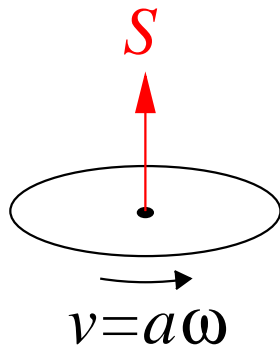
$$\vec{r}_{\text{CM}} \times \vec{p}_{\text{CM}} = (\vec{J} - \vec{S})_{\perp}$$

$$\gamma\beta mc(-x_{\text{CM}} \hat{y} + y_{\text{CM}} \hat{x}) = (\gamma - 1) \vec{S}_{\perp}^{\diamond}$$

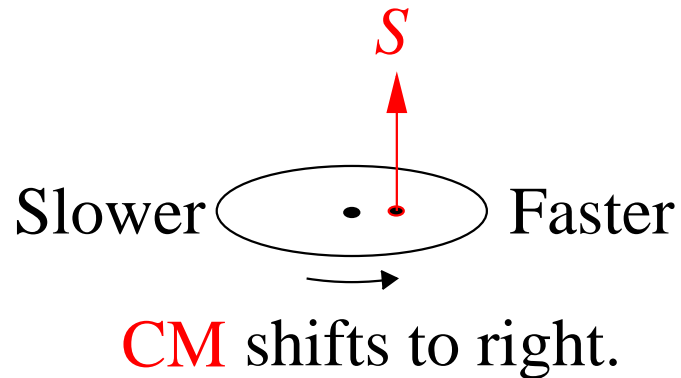
$$\gamma\beta mc(\vec{x}_{\text{CM}} + \vec{y}_{\text{CM}}) = (\gamma - 1) \hat{z} \times \vec{S}_{\perp}^{\diamond}$$

$$\vec{r}_{\perp \text{CM}} = \frac{\gamma}{\gamma + 1} \frac{\vec{\beta} \times \vec{S}}{mc}$$

CM at rest.



Boost into screen



Center of charge wobbles: classical “Zitterbewegung”

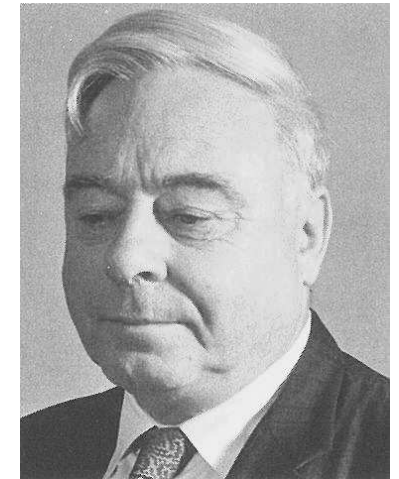


# Thomas Precession

1. Boost observer to left.
  2. Boost observer downward.
  3. Boost back to rest.
- Net rotation of rest frame.

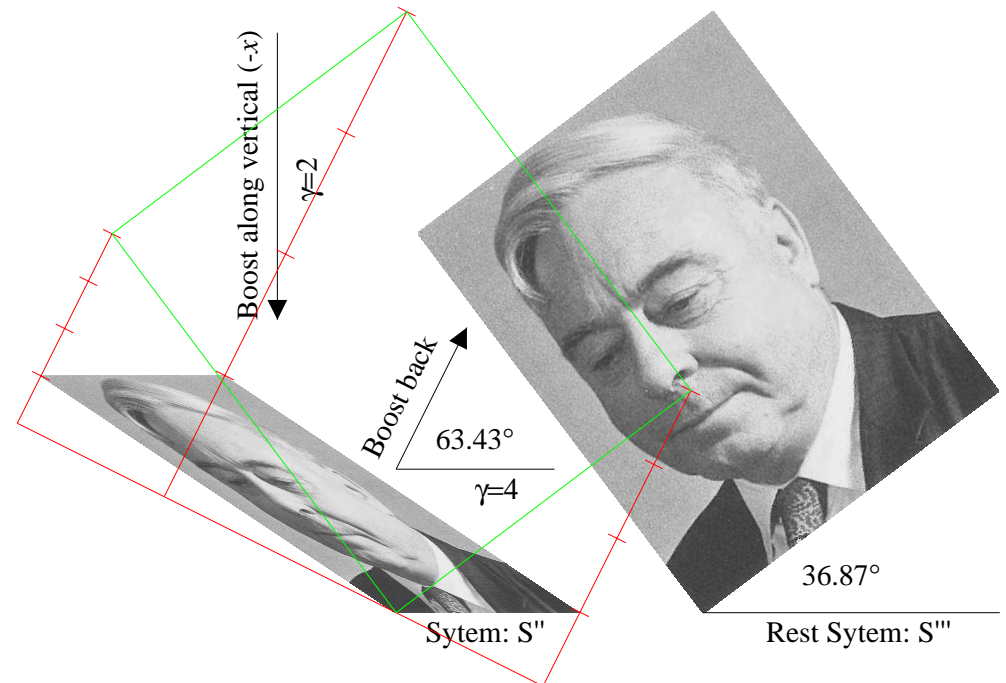


System: S'



Rest System: S

Boost along horizontal (-y)  
←  
 $\gamma=2$





# Thomas—Frenkel (BMT) Equation

In the local rest frame of the proton, the spin precession of the proton obeys the Thomas-Frenkel equation:

$$\frac{d\vec{S}^\diamond}{dt} = \frac{q}{\gamma m} \vec{S}^\diamond \times \left[ (1 + G\gamma)\vec{B}_\perp + (1 + G)\vec{B}_\parallel + \left( G\gamma + \frac{\gamma}{\gamma + 1} \right) \frac{\vec{E} \times \vec{v}}{c^2} \right].$$

This is a mixed description:  $t$ ,  $\vec{B}$ , and  $\vec{E}$  in the lab frame, but spin  $\vec{S}^\diamond$  in local rest frame of the particle:

Proton:  $G = \frac{g - 2}{2} = 1.792847,$  523.34 MeV/unit  $G\gamma$

Electron:  $a = G = \frac{g - 2}{2} = 0.001159652,$  440.65 MeV/unit  $a\gamma$

$$\gamma = \frac{\text{Energy}}{mc^2}.$$



# Thomas—Frenkel (BMT) Equation

In the local rest frame of the proton, the spin precession of the proton obeys the Thomas-Frenkel equation:

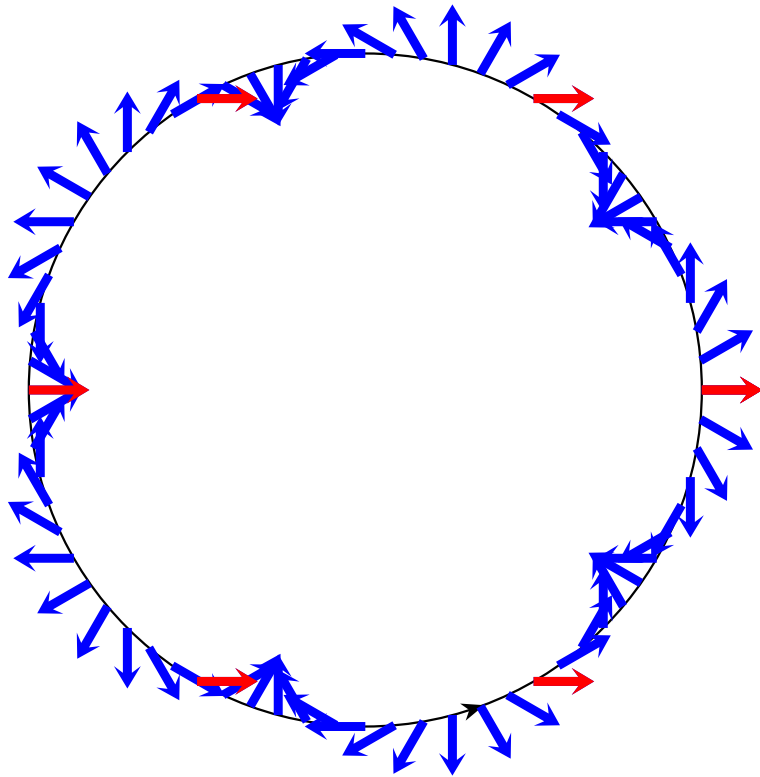
$$\begin{aligned} \text{Torque : } \quad \frac{d\vec{S}^\diamond}{dt} &= \frac{q}{\gamma m} \vec{S}^\diamond \times \left[ (1 + G\gamma)\vec{B}_\perp + (1 + G)\vec{B}_\parallel \right] && \text{TF} \\ \text{Force : } \quad \frac{d\vec{p}}{dt} &= \frac{q}{\gamma m} \vec{p} \times \vec{B}_\perp && \text{Lorentz} \end{aligned}$$

(This is a mixed description:  $t$ , and  $\vec{B}$  in the lab frame, but spin  $\vec{S}^\diamond$  in local rest frame of the proton.)

$$G = \frac{g - 2}{2} = 1.7928, \quad \gamma = \frac{\text{Energy}}{mc^2}.$$



# Spin Precession in a Ring



Example with 6 precessions of spin in one turn:

$$G\gamma + 1 = 6.$$

Spin tune: number of precessions per turn relative to beam's direction.

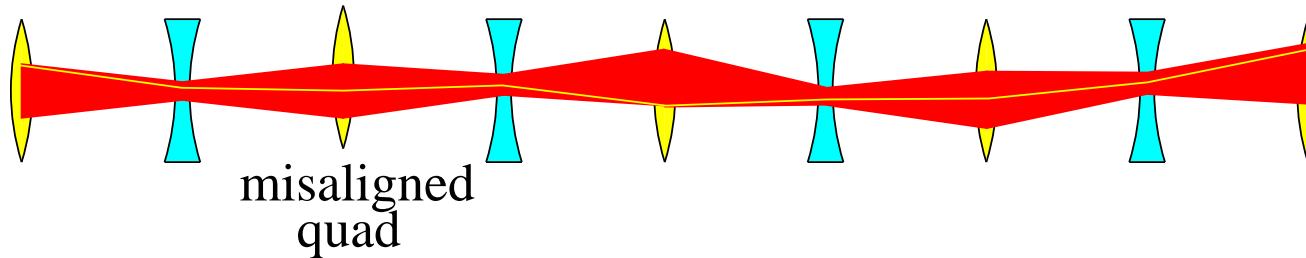
So we subtract one:

$$\nu_{\text{spin}} = G\gamma \propto \text{energy},$$

i.e., 5 in this example.



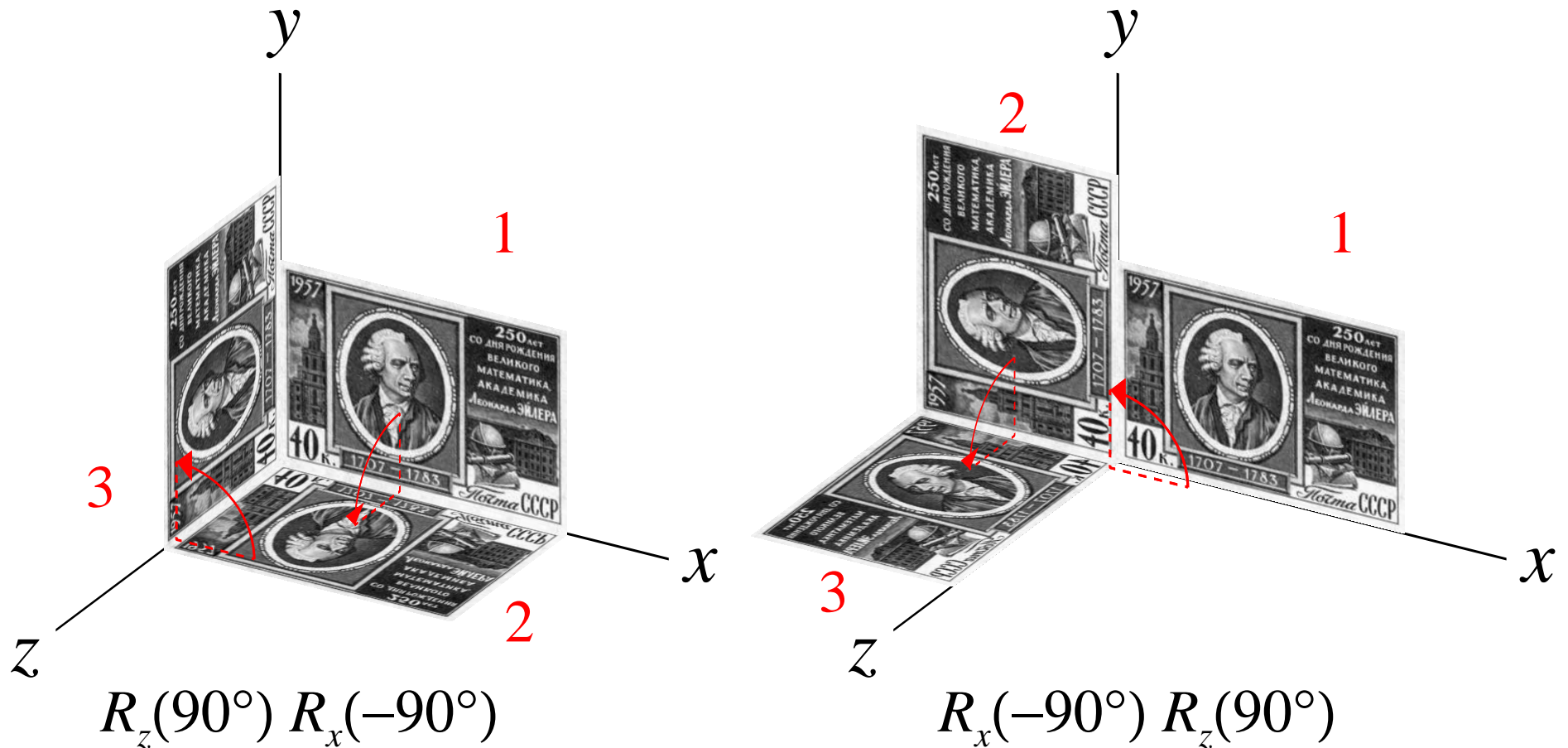
# Misalignments or Imperfections



- A misaligned quadrupole creates a steering error which propagates through the lattice.
- For an accelerator ring, this shifts the closed orbit away from the design trajectory.
- If the misalignment is vertical, then the design trajectory will have a periodic set of small vertical bends interspersed with the normal horizontal bends of the bending magnets.
- This leads to an integer resonance condition for the spin tune.



# In general, rotations don't commute.



# Hamiltonian with Spin

(Here I drop the “ $\diamond$ ” superscript on  $\vec{S}$ .)

$$\begin{aligned}\frac{d\vec{S}}{dt} &= \vec{W} \times \vec{S} \\ H(\vec{q}, \vec{P}, \vec{S}; s) &= \mathcal{H}_{\text{orb}} + \mathcal{H}_{\text{spin}} \\ &= \mathcal{H}_{\text{orb}} + \vec{W} \cdot \vec{S} + O(\hbar^2)\end{aligned}$$

Group symmetries:

- Orbital motion without spin:  $\text{Sp}(6, r)$ .
- Spin by itself:  $\text{SU}(2, c) \cong \text{SO}(3, r)$  (homomorphic).
- Full blown symmetry:  $\text{Sp}(6, r) \oplus \text{SU}(2, c)$ .
  - Spin dependence on orbit (Thomas-Frenkel).
  - Orbit dependence on spin (Stern-Gerlach Force)—Usually ignored.



The differential equation for the spin vector:

$$\frac{d\vec{S}}{dt} = \vec{W} \times \vec{S}$$

can be written as the matrix equation:

$$\frac{d}{dt} \begin{pmatrix} S_x \\ S_y \\ S_z \end{pmatrix} = \begin{pmatrix} 0 & -W_z & W_y \\ W_z & 0 & -W_x \\ -W_y & W_x & 0 \end{pmatrix} \begin{pmatrix} S_x \\ S_y \\ S_z \end{pmatrix}.$$

This equation may be integrated via Lie algebraic techniques as was done in Chapter 3 of Conte & MacKay.

We can calculate spin rotations matrices for a rotation of spin vector by an angle  $\theta$  about a vector  $\hat{n}$ .



# Representation of Rotations

SO(3) :

$$\mathbf{R}_{\hat{n}}(\theta) = \mathbf{I} \cos \theta + \begin{pmatrix} 0 & n_z & -n_y \\ -n_z & 0 & n_x \\ n_y & -n_x & 0 \end{pmatrix} \sin \theta + \begin{pmatrix} n_x^2 & n_x n_y & n_x n_z \\ n_x n_y & n_y^2 & n_y n_z \\ n_x n_z & n_y n_z & n_z^2 \end{pmatrix} (1 - \cos \theta).$$

SU(2) with usual spinor notation:

$$\text{Pauli matrices: } \sigma_x = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \sigma_y = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}, \sigma_z = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}.$$

$$\mathbf{R}_{\hat{n}}(\theta) = e^{i \hat{n} \cdot \vec{\sigma} \theta / 2} = \begin{pmatrix} \cos \frac{\theta}{2} + i n_z \sin \frac{\theta}{2} & (n_y + i n_x) \sin \frac{\theta}{2} \\ (-n_y + i n_x) \sin \frac{\theta}{2} & \cos \frac{\theta}{2} - i n_z \sin \frac{\theta}{2} \end{pmatrix}.$$





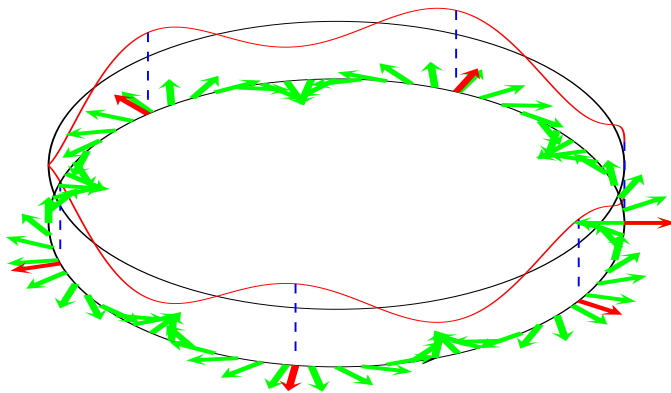
# Depolarizing Resonances

Simple Resonance Condition:

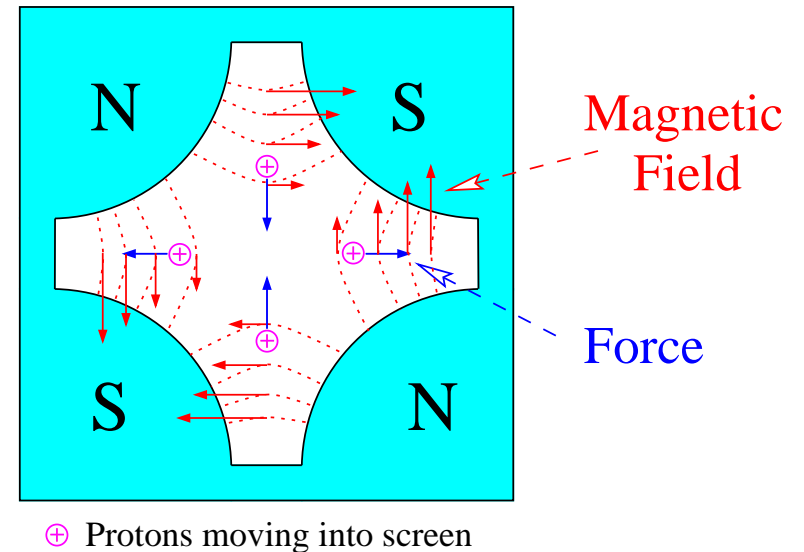
$$\nu_{\text{spin}} = N + N_v Q_v,$$

(imperfection)                      (intrinsic)

where  $N$  and  $N_v$  are integers.



**Magnetic Lens (quadrupole)**  
Vertically focusing



# Example: Ring with one partial snake

1-turn matrix:

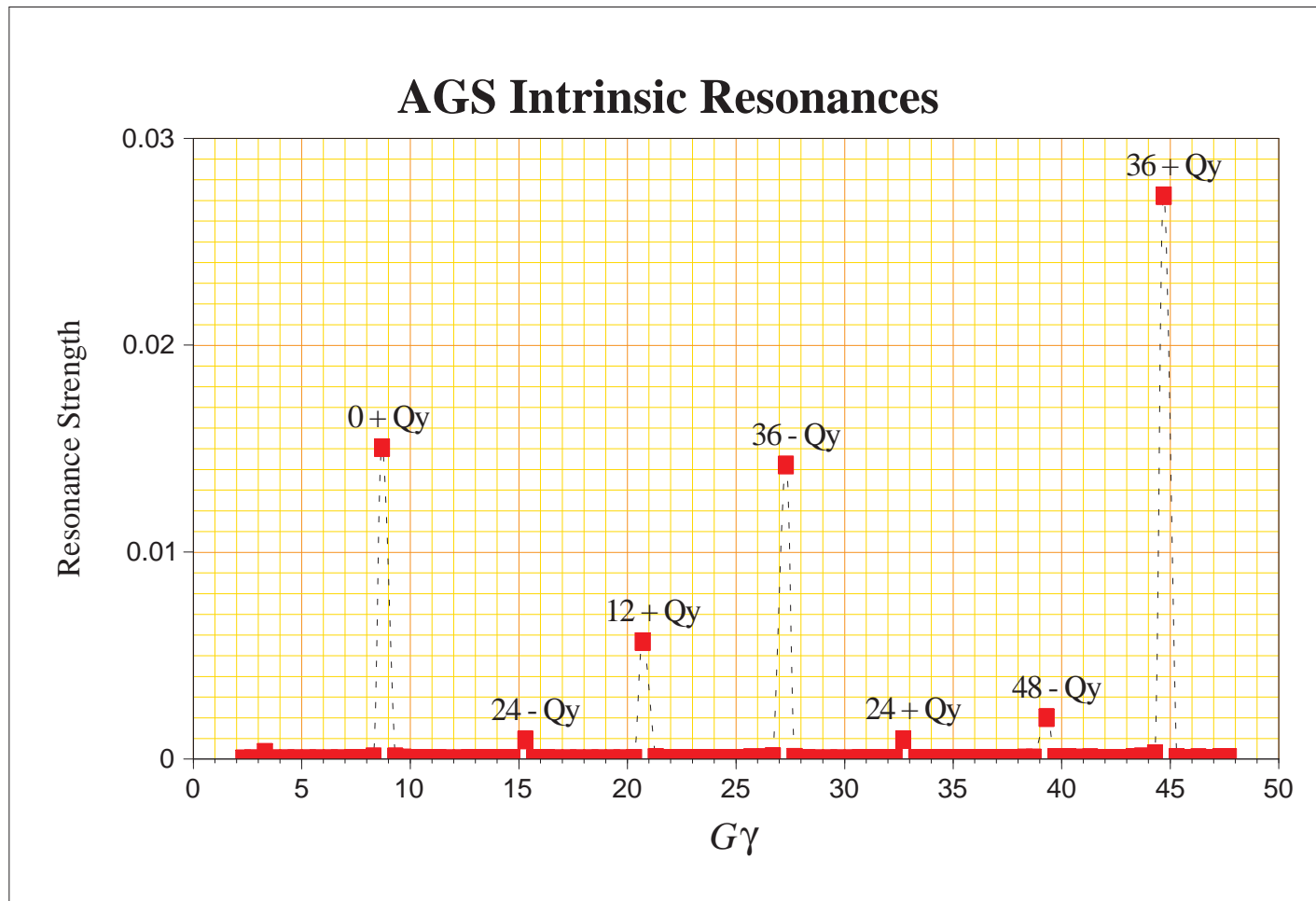
$$\begin{aligned}\mathbf{M} &= \begin{pmatrix} \cos \frac{\theta}{2} + in_z \sin \frac{\theta}{2} & (n_y + in_x) \sin \frac{\theta}{2} \\ (-n_y + in_x) \sin \frac{\theta}{2} & \cos \frac{\theta}{2} - in_z \sin \frac{\theta}{2} \end{pmatrix} \\ &= \mathbf{R}_{\hat{y}}(G\gamma\pi) \mathbf{R}_{\hat{z}}(\mu/2) \\ &= \begin{pmatrix} \cos(G\gamma\pi) & \sin(G\gamma\pi) \\ -\sin(G\gamma\pi) & \cos(G\gamma\pi) \end{pmatrix} \begin{pmatrix} \exp(i\mu/2) & 0 \\ 0 & \exp(-i\mu/2) \end{pmatrix} \\ &= \begin{pmatrix} \exp(i\mu/2) \cos(G\gamma\pi) & \exp(-i\mu/2) \sin(G\gamma\pi) \\ -\exp(i\mu/2) \sin(G\gamma\pi) & \exp(-i\mu/2) \cos(G\gamma\pi) \end{pmatrix} \\ \\ \frac{1}{2} \text{tr}(\mathbf{M}) &= \cos \frac{\mu}{2} \cos(G\gamma\pi) = \cos \frac{\theta}{2} = \cos(\pi\nu_{\text{spin}})\end{aligned}$$

since  $\theta = 2\pi\nu_{\text{spin}}$ .

A snake is a device which rotates the spin about a vector in the horizontal plane.



# AGS Intrinsic Resonances



Without snakes



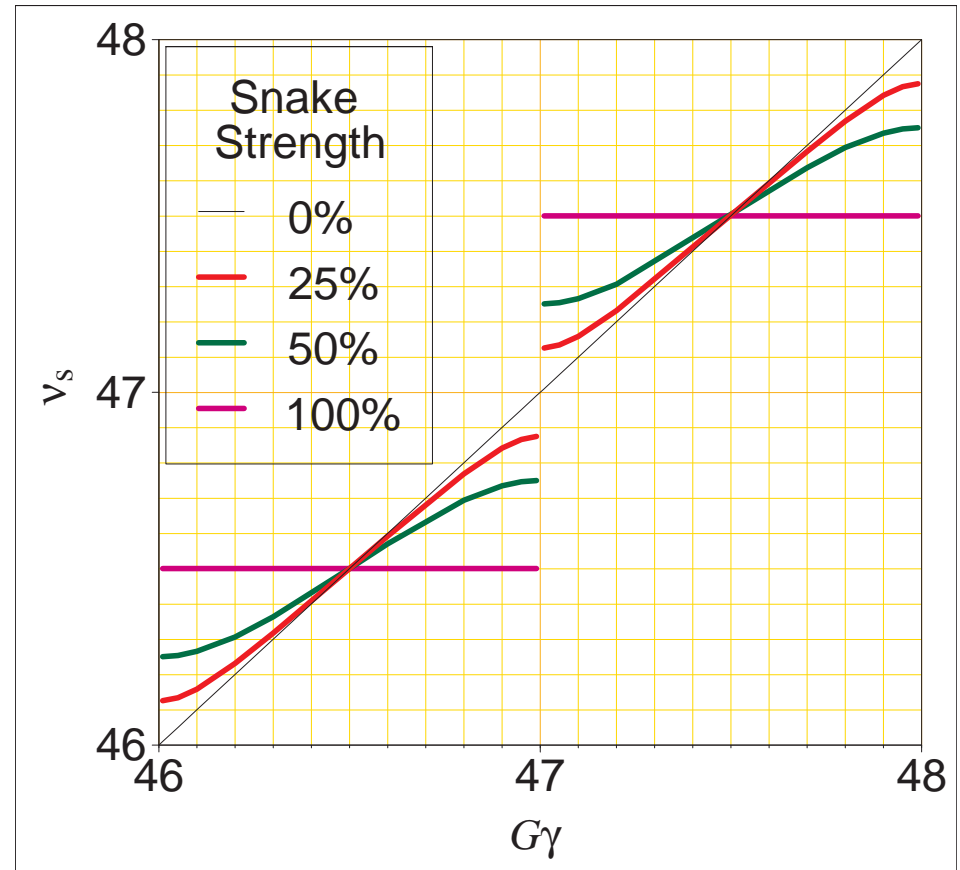
# Partial Snakes

Adding a partial snake opens up stop bands around the integer imperfection resonances.

At the snake the stable spin direction points along the snake's rotation axis when  $G\gamma = \text{integer}$ .

Partial snake strength:  $\frac{\mu}{\pi}$

$$\cos \pi \nu_s = \cos(G\gamma\pi) \cos \frac{\mu}{2}$$



# 🎪 Crossing an Isolated Spin Resonance 🎪

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Froissart—Stora Formula:

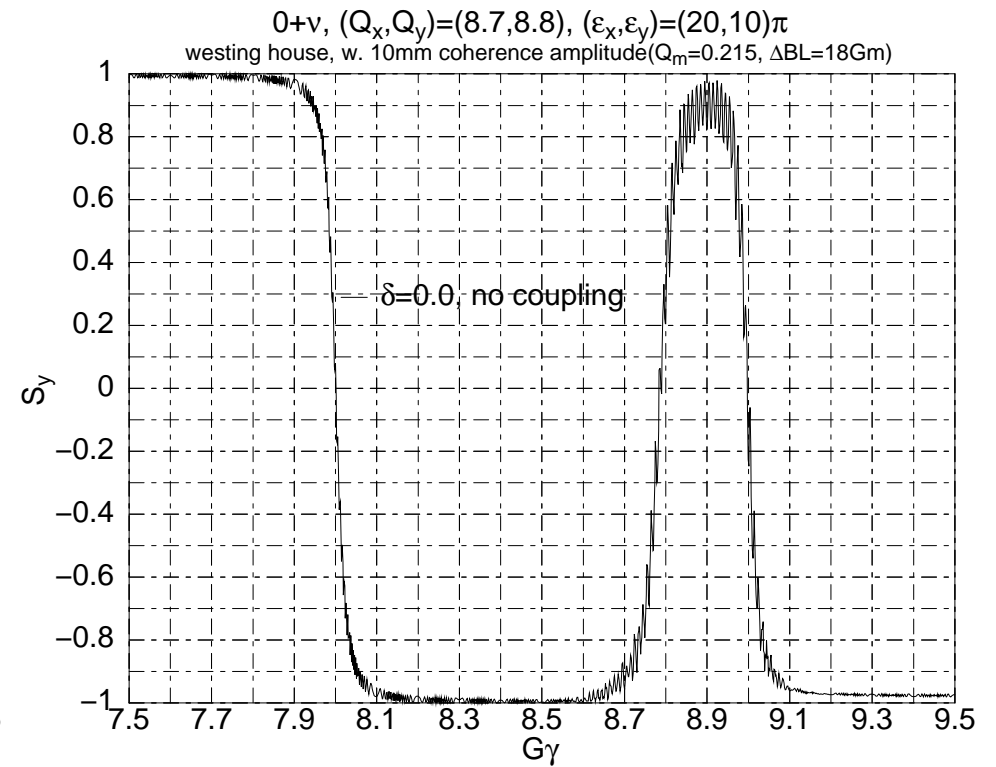
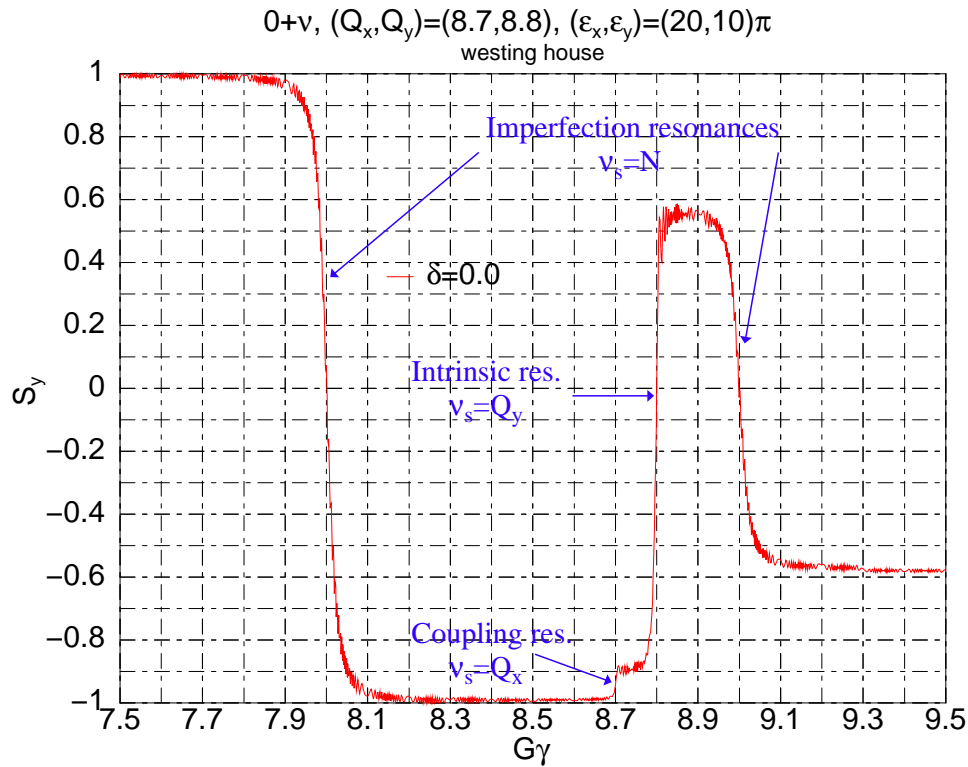
$$\frac{P_f}{P_i} = 2 \exp\left(-\frac{\pi|\epsilon|^2}{2\alpha}\right) - 1.$$

Ramp rate:  $\alpha = \frac{dG\gamma}{d\theta}$ , ( $\theta : 2\pi/\text{turn.}$ )

Resonance strength:  $\epsilon = \text{Fourier amplitude.}$



# Resonance Crossing in AGS

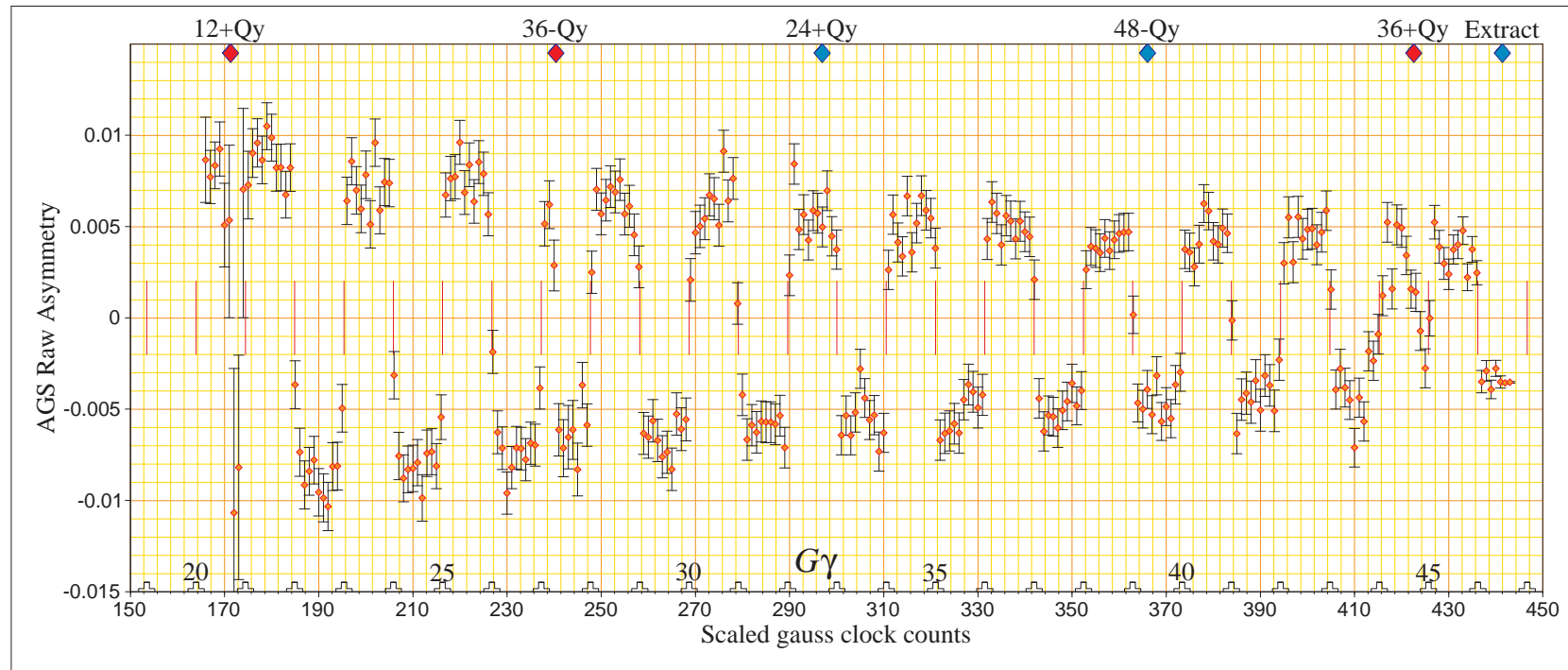


AC dipole used to increase strength of  $\nu_{\text{spin}} = Q_y$  resonance.

(Simulations by Mei Bai)



# AGS Raw Asymmetry during Ramp



AGS has 12 superperiods.  
Vertical betatron tune: 8.7  
Snake strength: 5%

(From Jeff Woods)

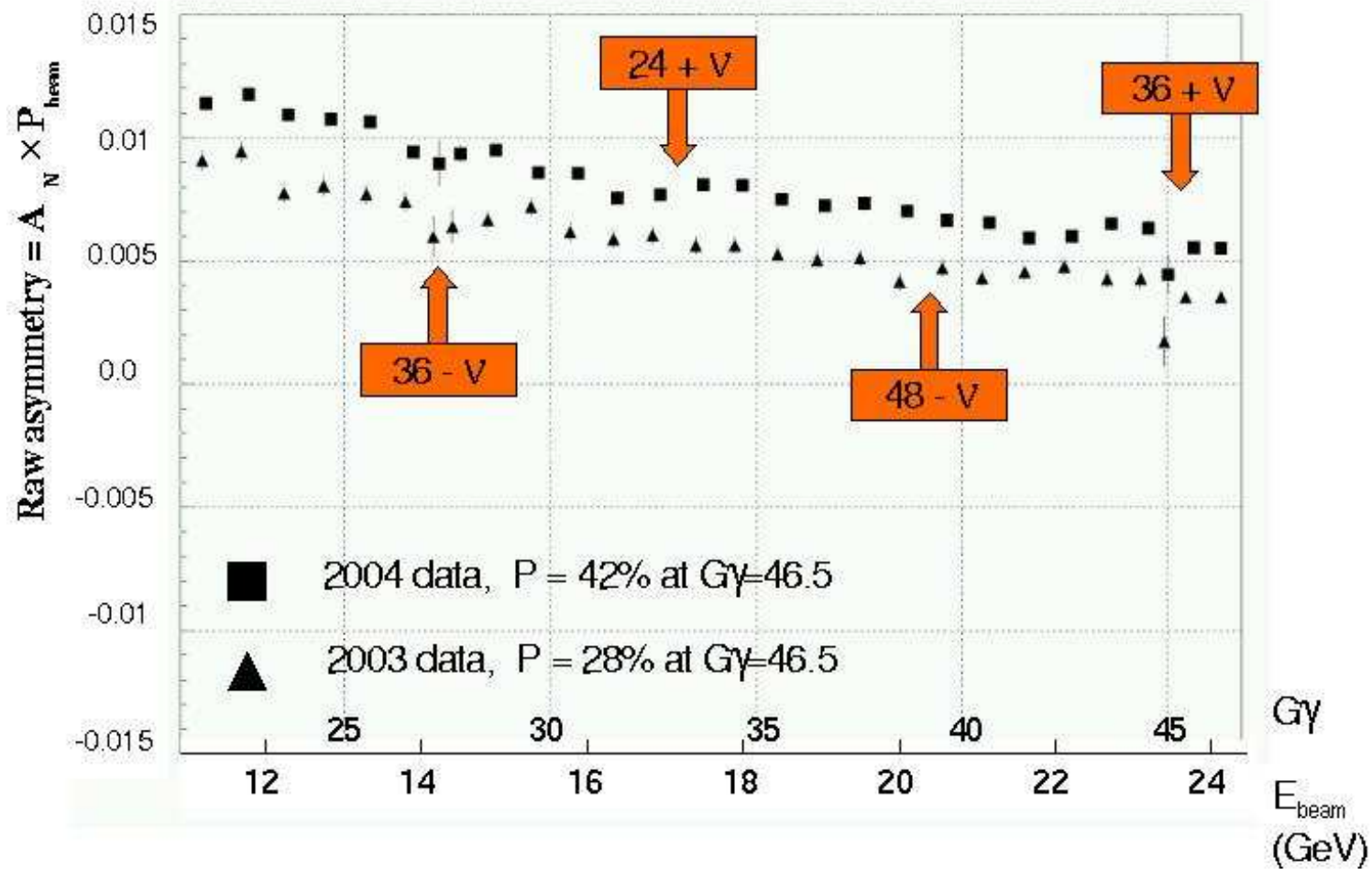
AC dipole pulses at resonances:

- $0 + Q_y$
- $12 + Q_y$
- $36 - Q_y$
- $36 + Q_y$

USPAS: Lecture on Spin Dynamics in Accelerators  
Waldo MacKay January, 2013

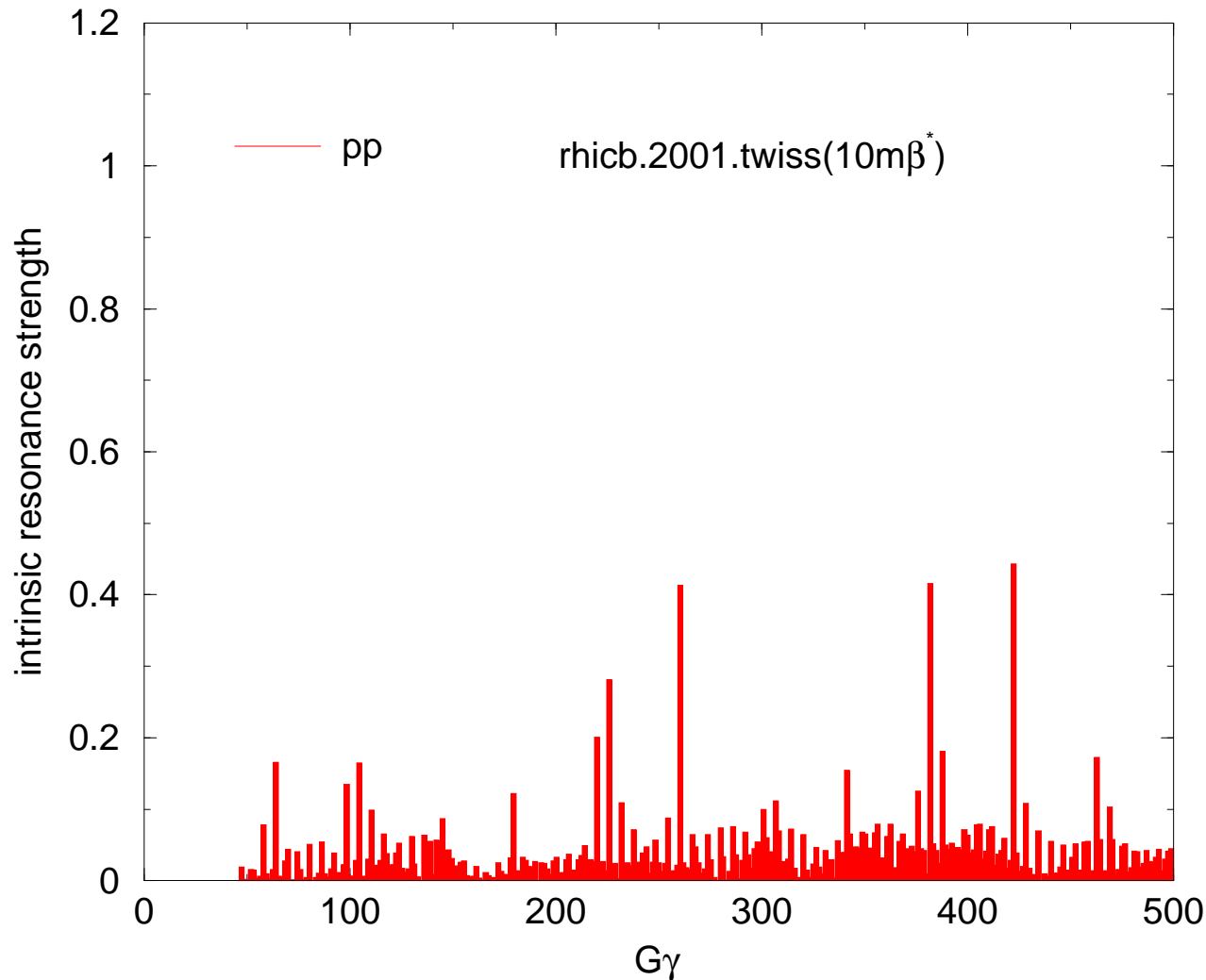


# pC CNI Asymmetry during AGS Ramp





# Depolarizing Resonances in RHIC



$$Q_x = 28.236$$

$$Q_y = 29.219$$

$$\pi\epsilon_y = 10\pi \mu\text{m}$$

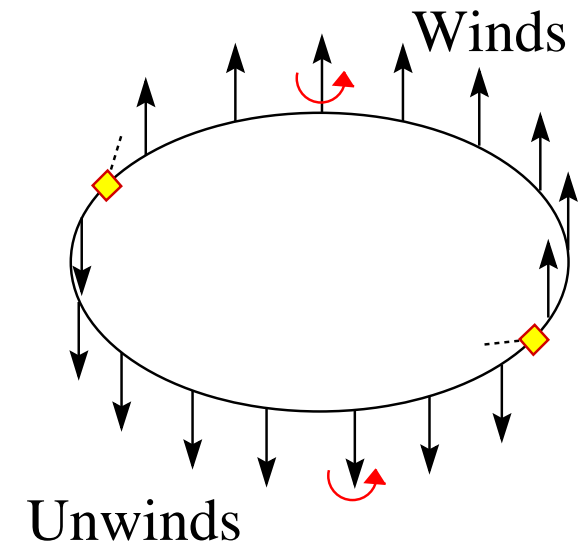
Will depolarize beam during acceleration.

Solution: Snakes

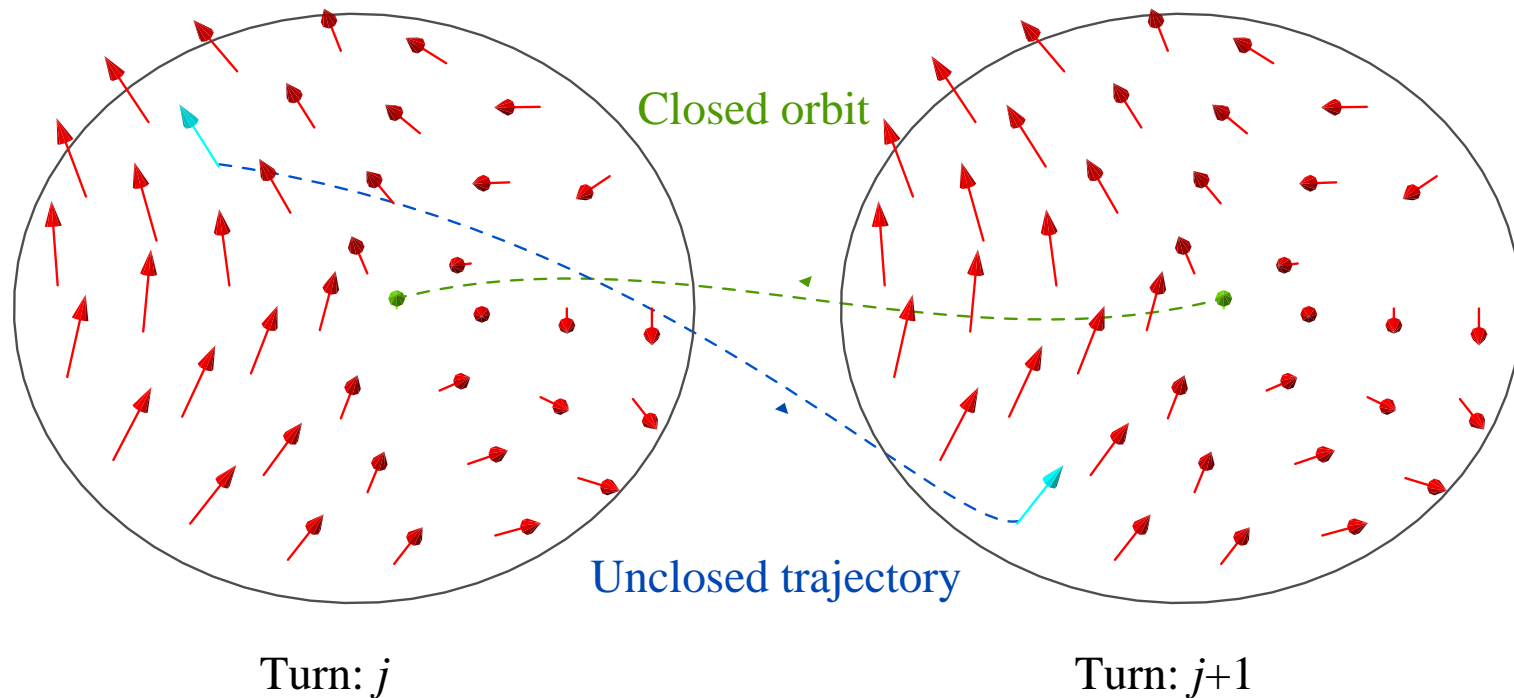


# Snake Charming

- 2 snakes: spin is up in one half of the ring, and down in the other half.
- Spin tune:  $\nu_{\text{spin}} = \frac{1}{2}$   
(It's energy independent.)
- “The unwanted precession which happens to the spin in one half of the ring is unwound in the other half.”



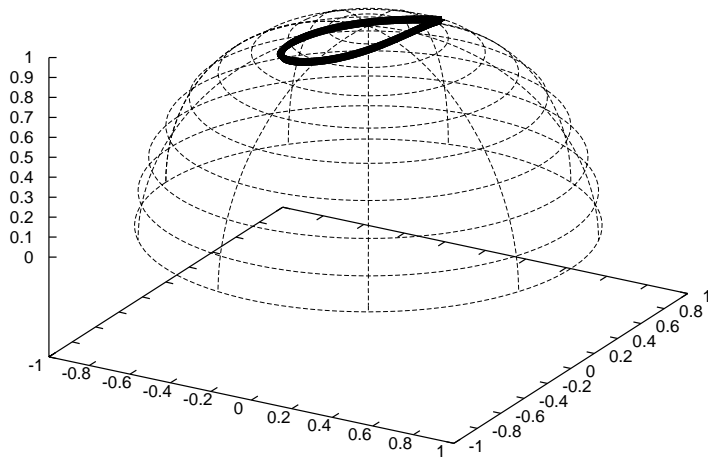
# Invariant Spin Field



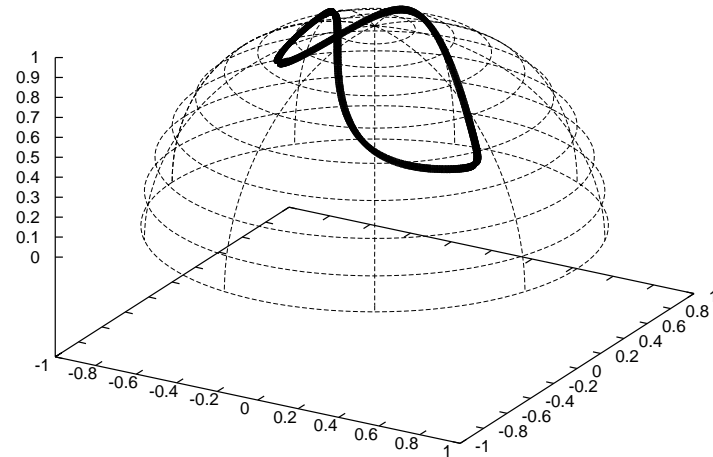
- For the closed orbit:  $\vec{n}_0(s) = \vec{n}_0(s + L)$ ,  
with  $\vec{q}_0(s) = \vec{q}_0(s + L)$  and  $\vec{P}_0(s) = \vec{P}_0(s + L)$ .
- For other locations in phase space:  $\vec{n}(\vec{q}, \vec{P}, s) = \vec{n}(\vec{q}, \vec{P}, s + L)$ ,  
even though in general  $q(s + L) \neq q(s)$  and  $P(s + L) \neq P(s)$ .

# HERA-p: Invariant Spin Field

a: HERA-p / 8 snakes / 4 pi mm mrad / 800 GeV



b: HERA-p / 8 snakes / 4 pi mm mrad / 802 GeV



$\hat{n}$ -vector at  $1\sigma$  and 800 GeV

$\hat{n}$ -vector at  $1\sigma$  and 802 GeV

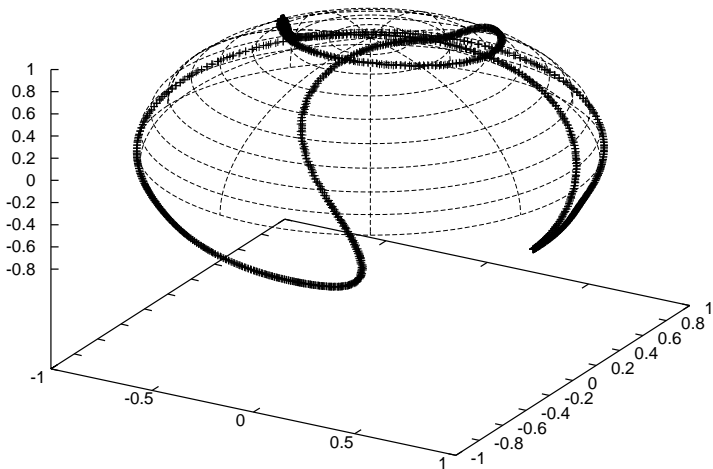
- Simulation with only vertical betatron motion.
- 802 GeV is closer to a resonance spin resonance than 800 GeV.

Des Barber et al.

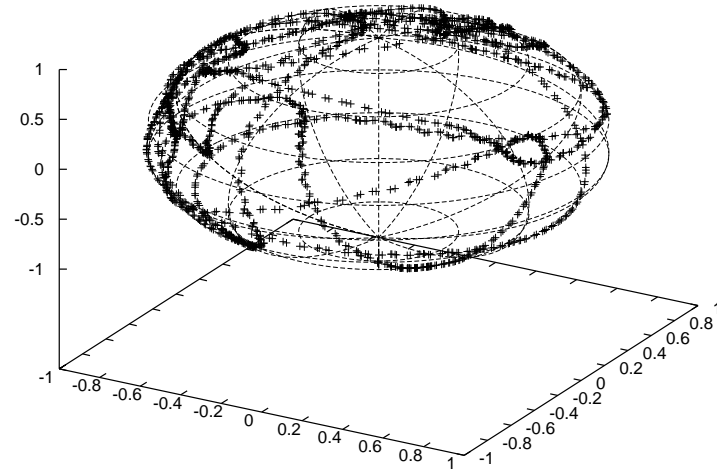


# HERA-p: Invariant Spin Field

a: HERA-p / 8 snakes / 64 pi mm mrad / 800 GeV



b: HERA-p / 8 snakes / 64 pi mm mrad / 802 GeV



$\hat{n}$ -vector at  $4\sigma$  and 800 GeV

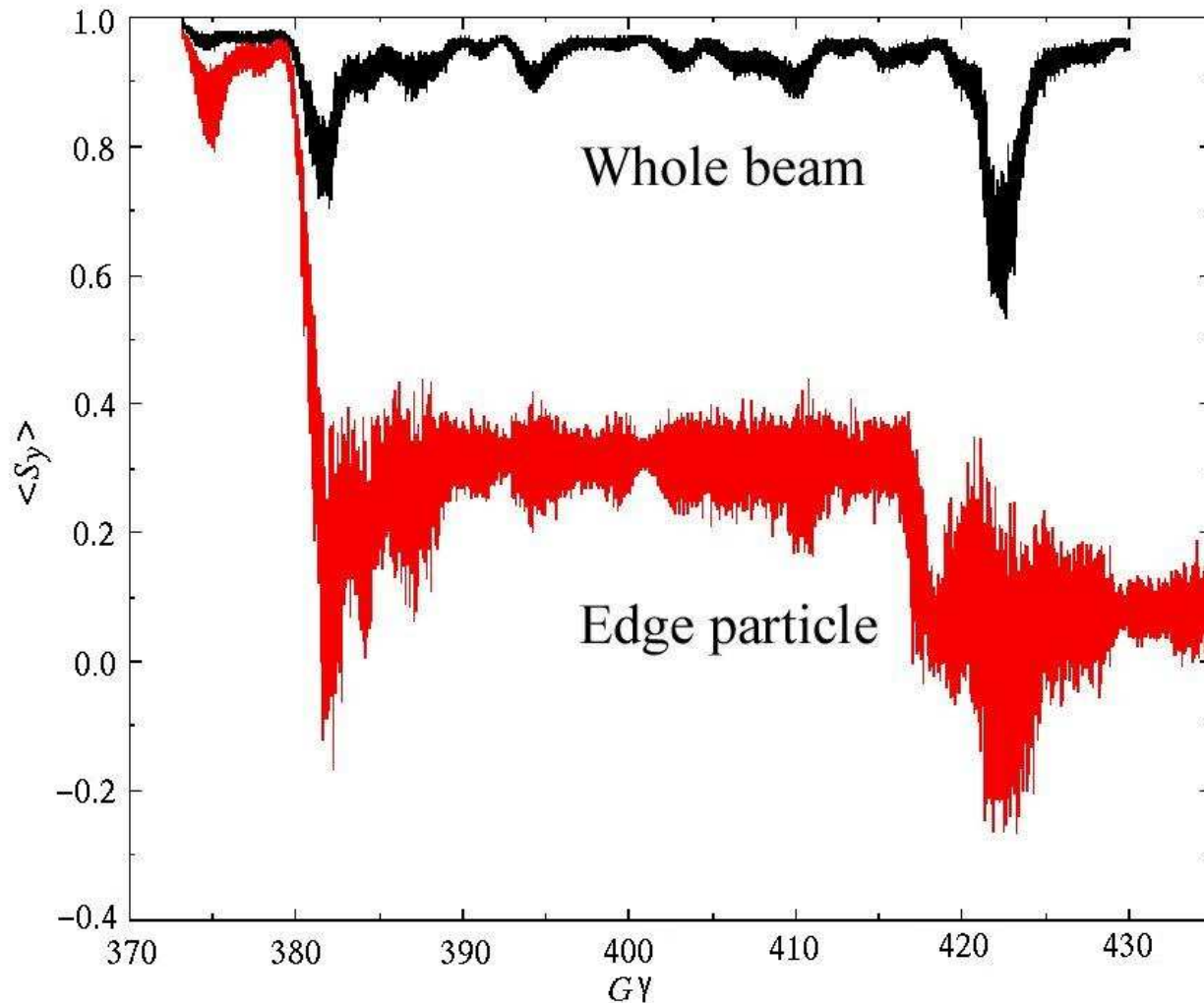
$\hat{n}$ -vector at  $4\sigma$  and 802 GeV

- Larger amplitude oscillations have a larger tune shift due to nonlinear elements.
- 802 GeV is closer to a resonance spin resonance than 800 GeV.

Des Barber et al.



# Spin Tracking in RHIC

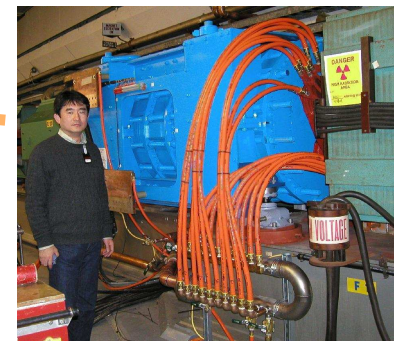
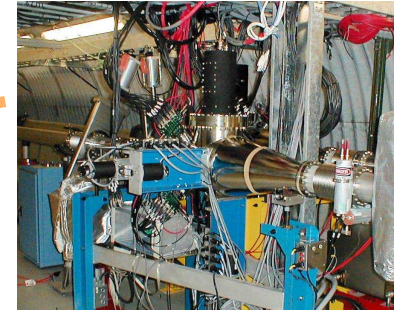
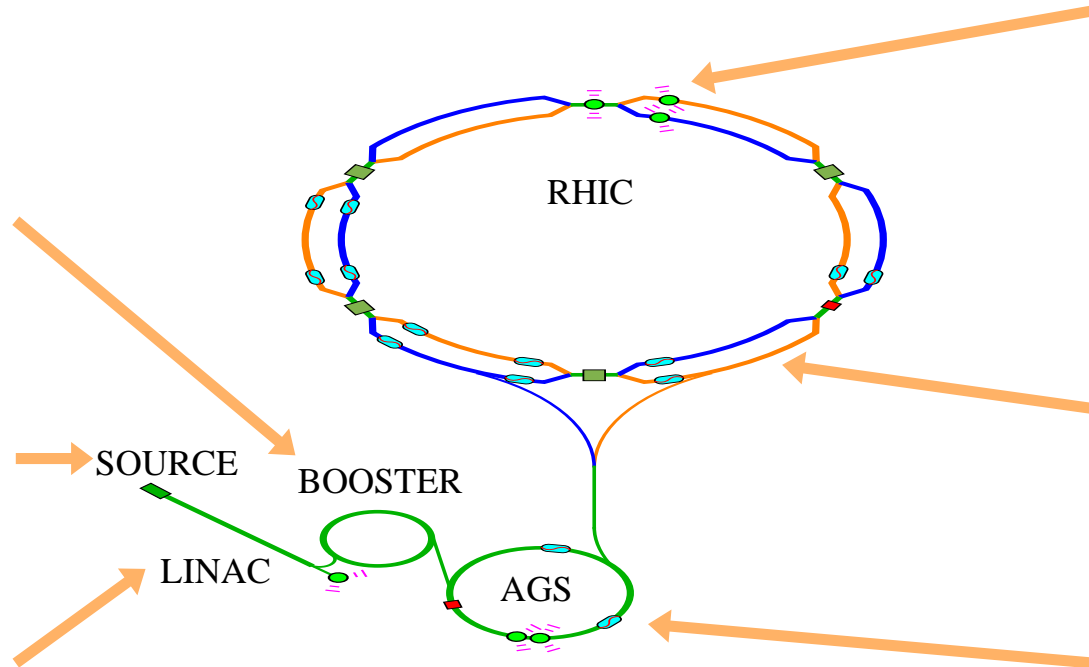
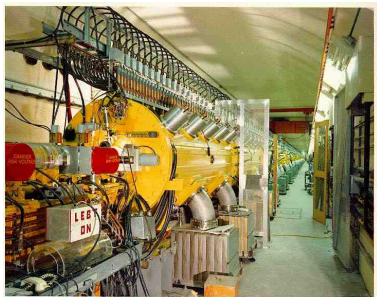
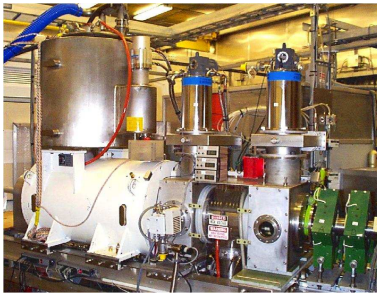


Particles with larger amplitude betatron oscillations may experience more precession away from the stable spin direction of the center of the beam

(Alfredo Luccio)



# Accelerators with Polarized Protons

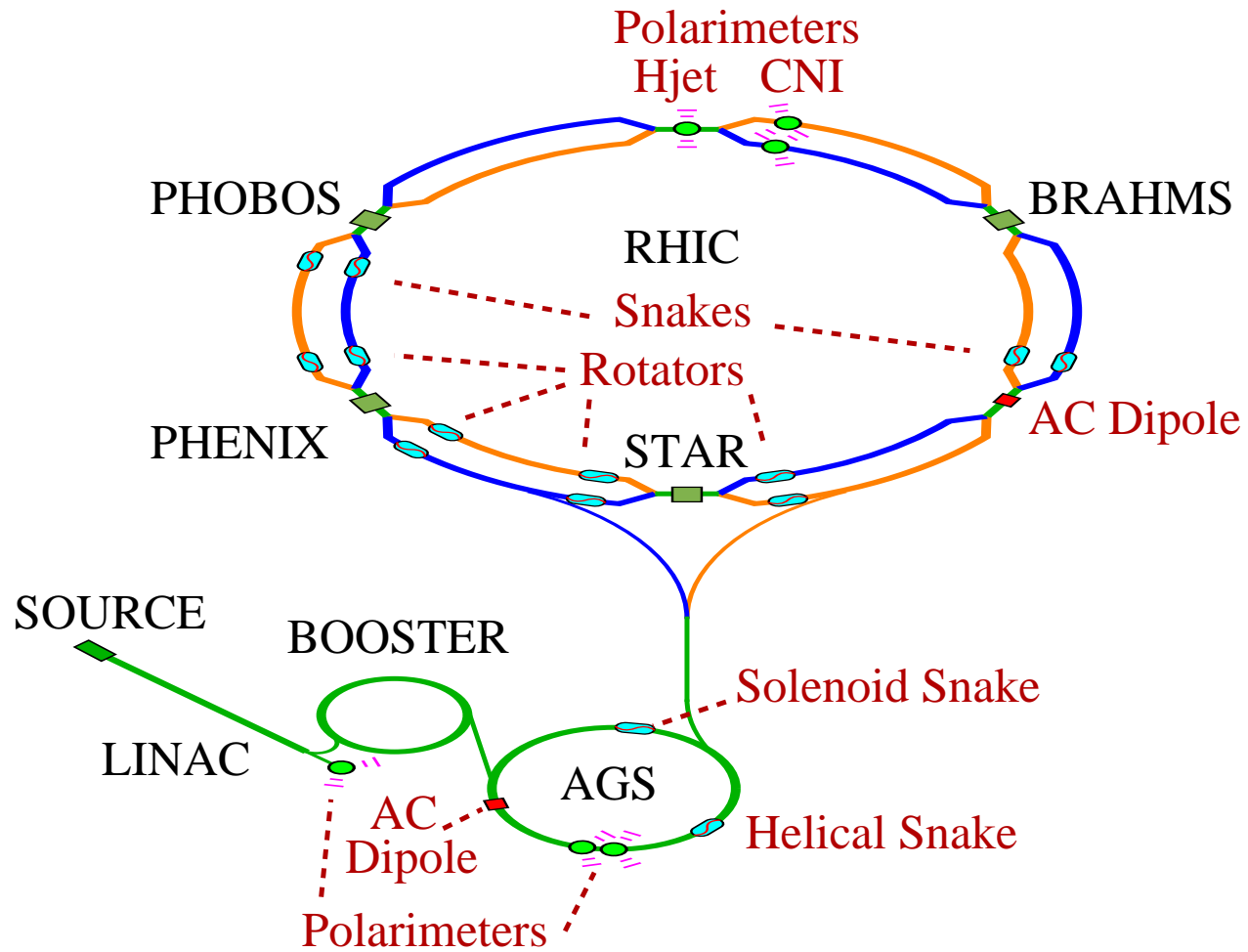


LINAC: Linear Accelerator  
AGS: Alternating Gradient Synchrotron  
RHIC: Relativistic Heavy Ion Collider



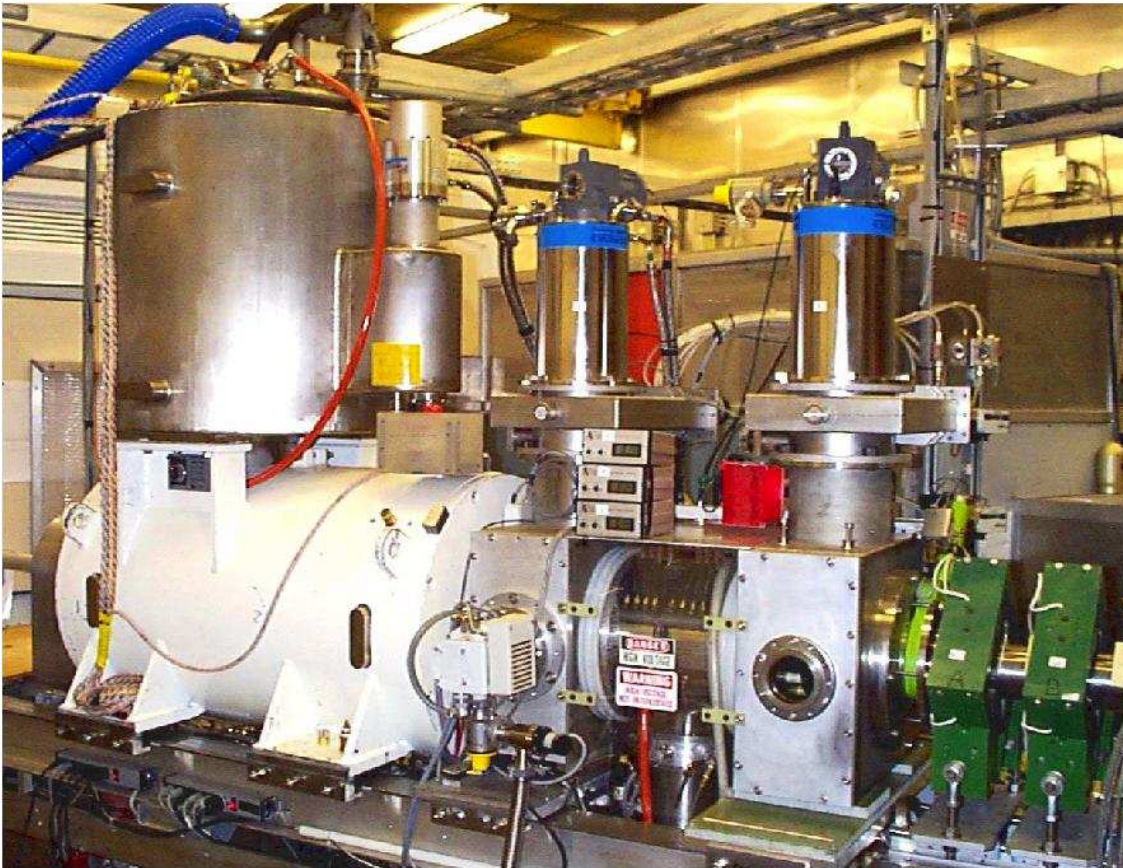


# Accelerator Complex for Protons





# High Intensity Polarized H<sup>-</sup> Source



KEK OPPIS\*  
upgraded at TRIUMF  
70 → 80% Polarization  
 $15 \times 10^{11}$  protons/pulse  
at source  
 $6 \times 10^{11}$  protons/pulse  
at end of LINAC

\*Optically Pumped Polarized Ion Source

# Optically Pumped Polarized Ion Source

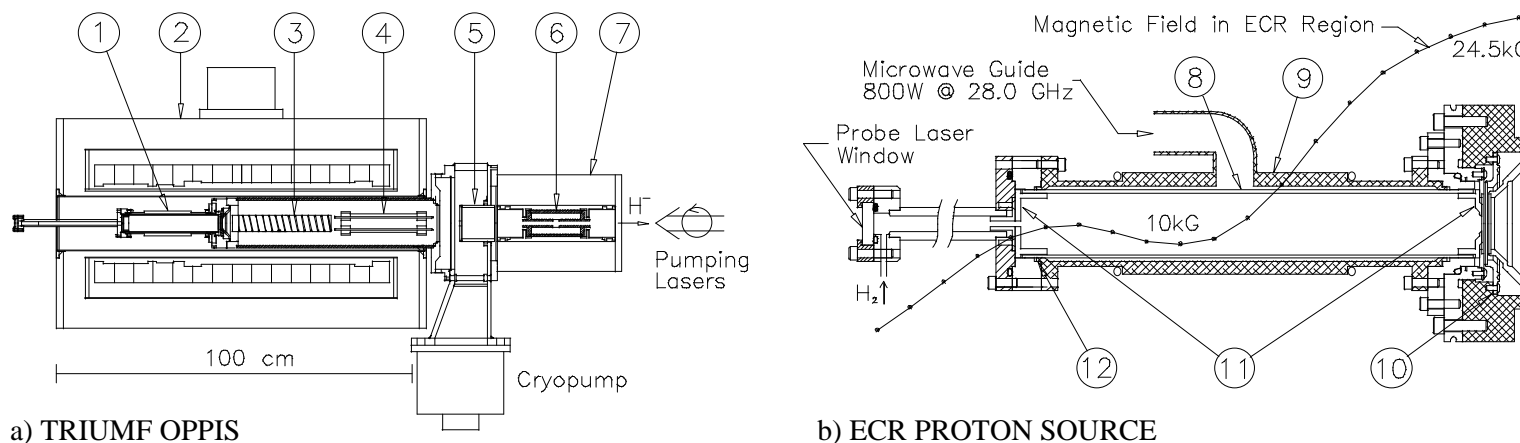
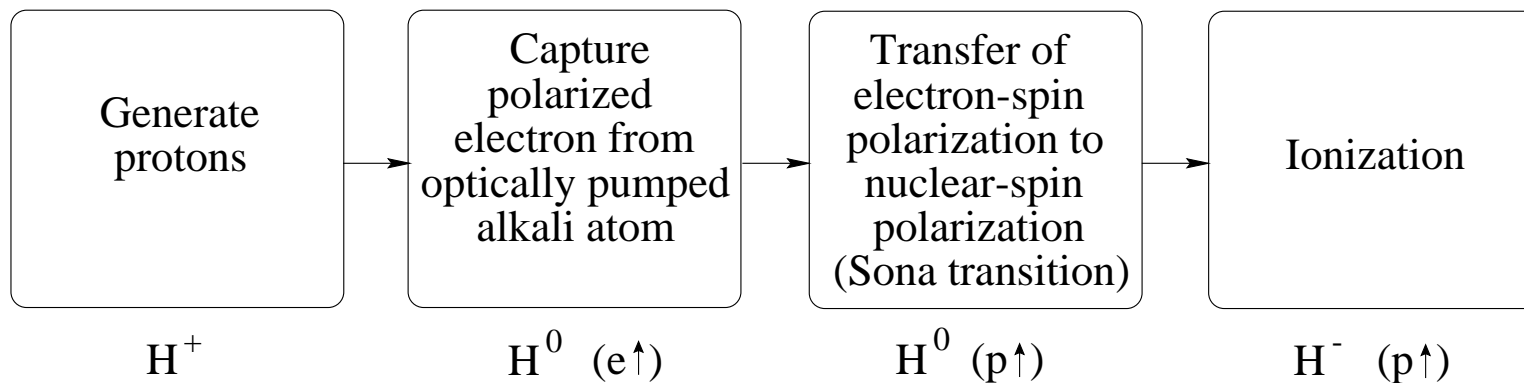
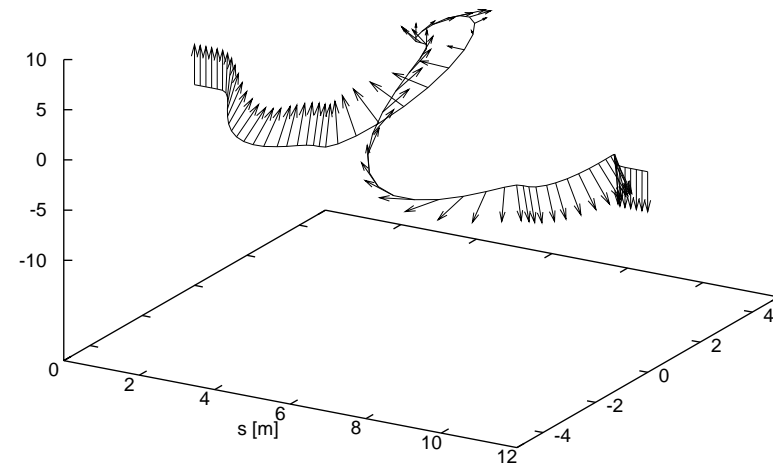
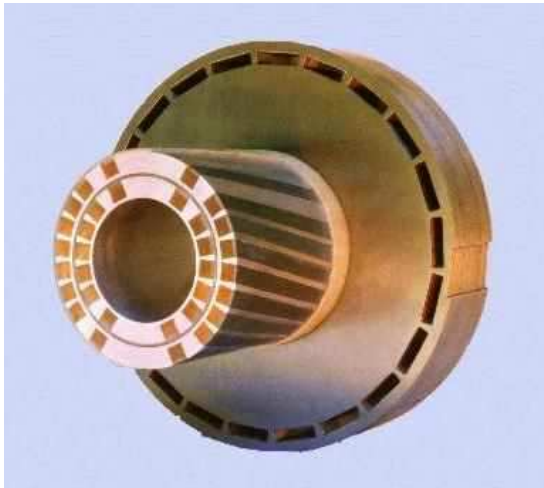
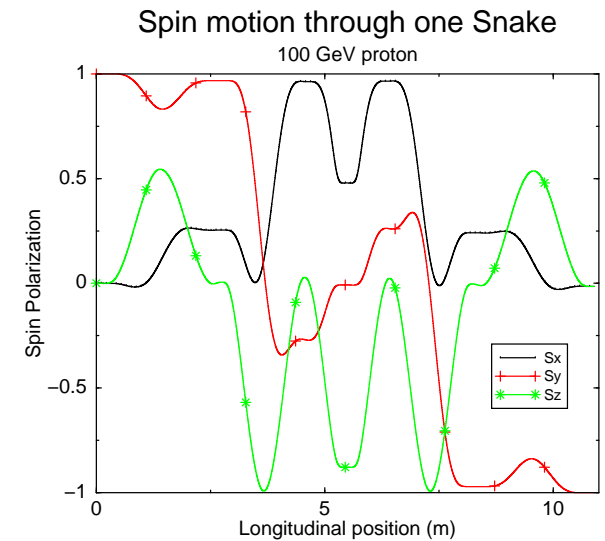
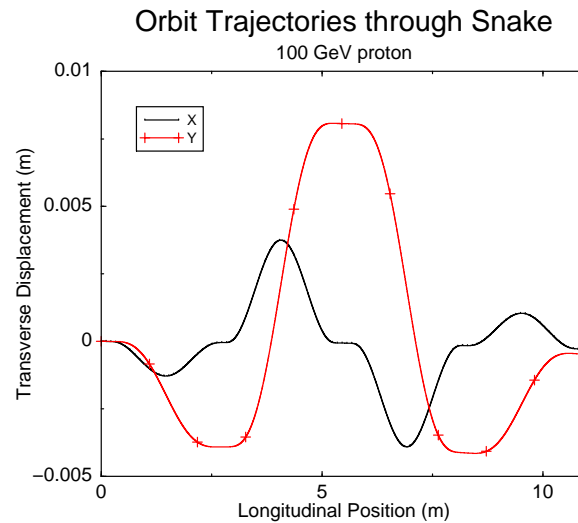
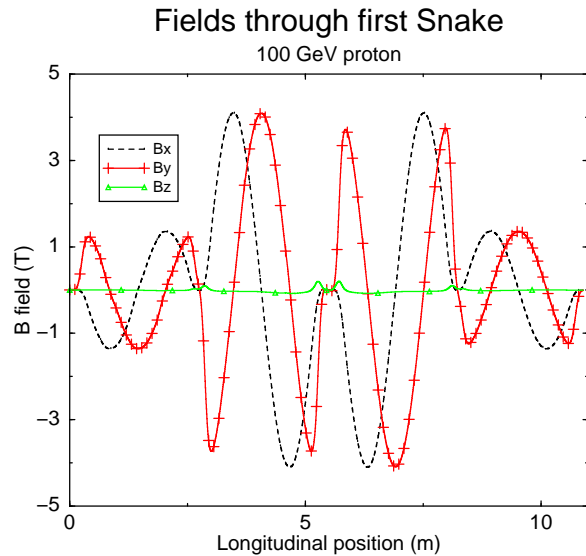


Fig. 1. 1) ECR Proton Source, 2) Superconducting Solenoid, 3) Optically-Pumped Rb Cell, 4) Deflection Plates, 5) Sona Transition Region, 6) Ionizer Cell, 7) Ionizer Solenoid, 8) Quartz Tube, 9) ECR Cavity, 10) Three Grid Extraction System, 11) Boron-Nitride End Cups, 12) Indium Seals.

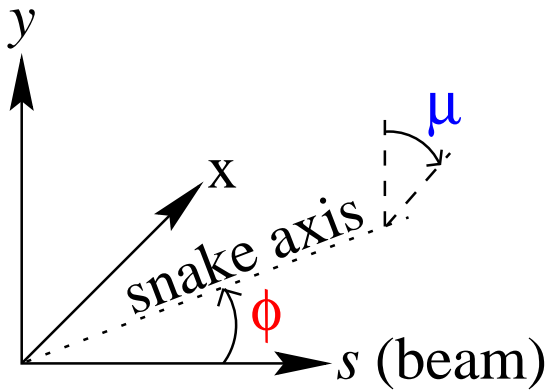


# Trajectory and Spin through Snakes

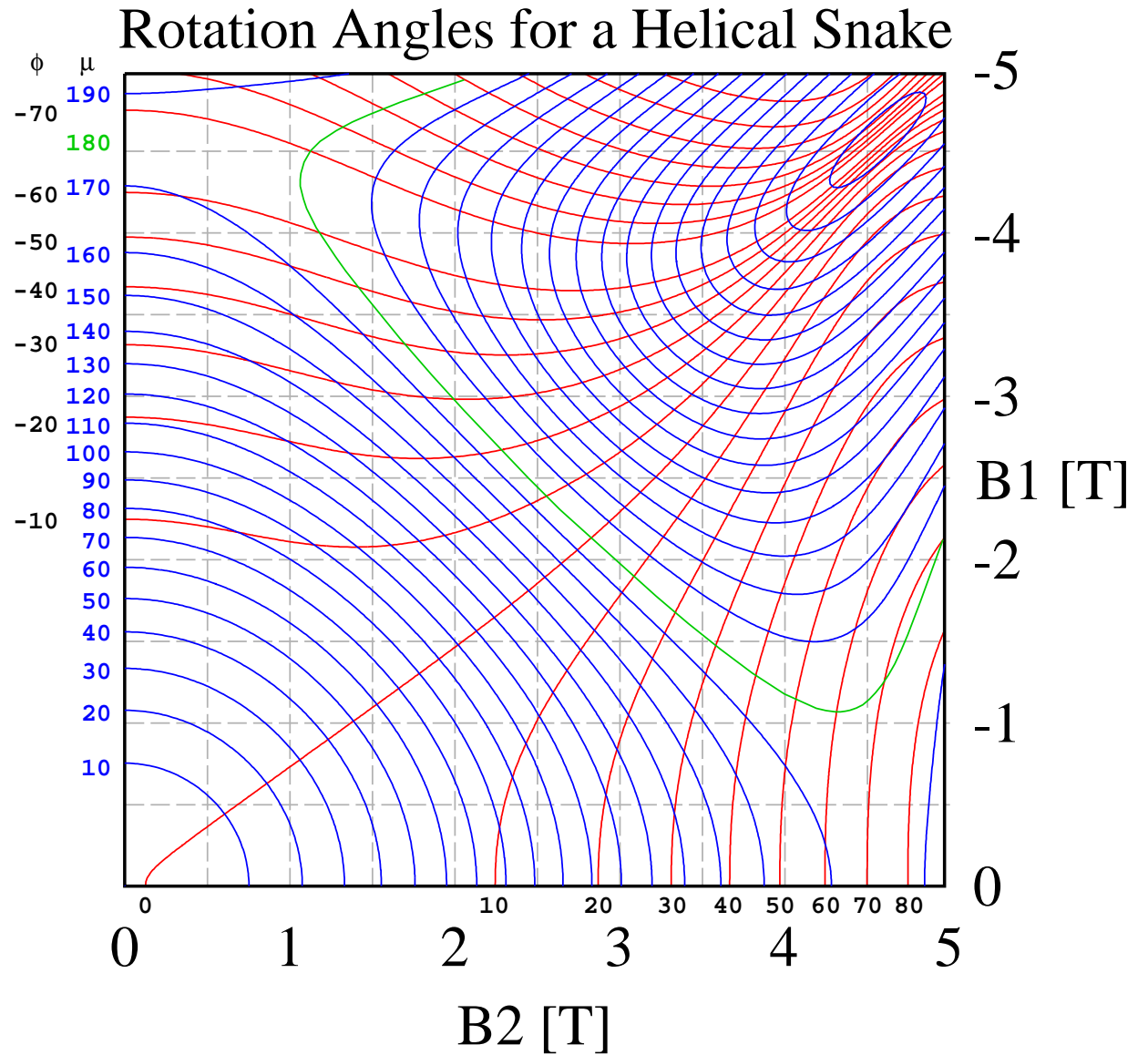


USPAS: Lecture on Spin Dynamics in Accelerators  
Waldo MacKay January, 2013

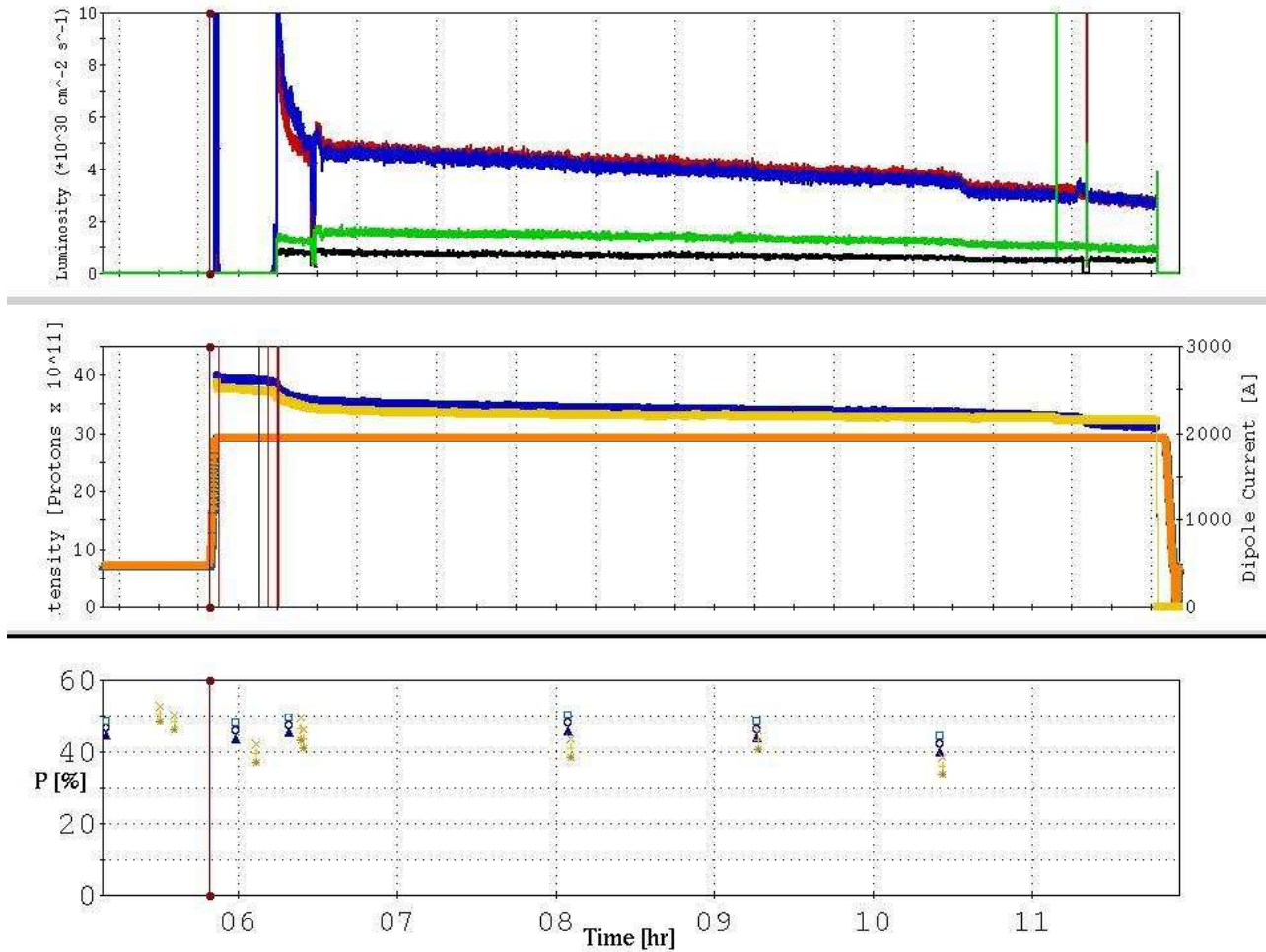




The rotation axis of the snake is  $\phi$ , and  $\mu$  is the rotation angle.



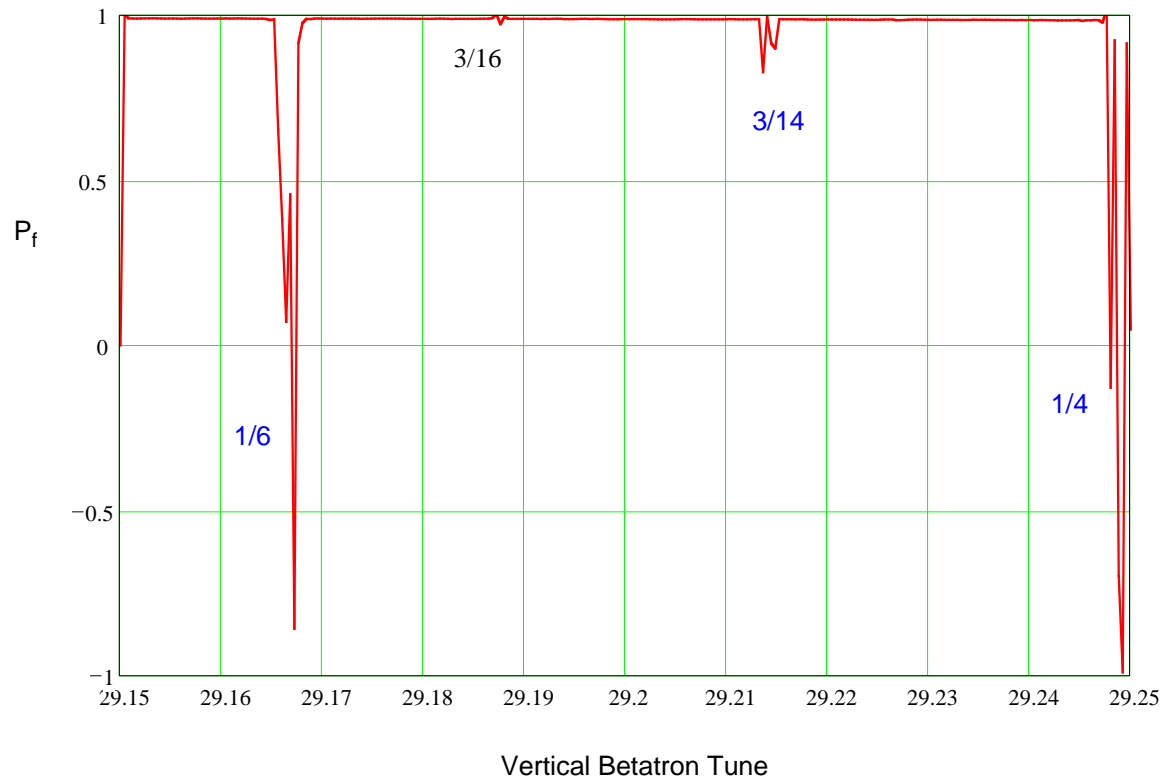
# RHIC Beam Polarization



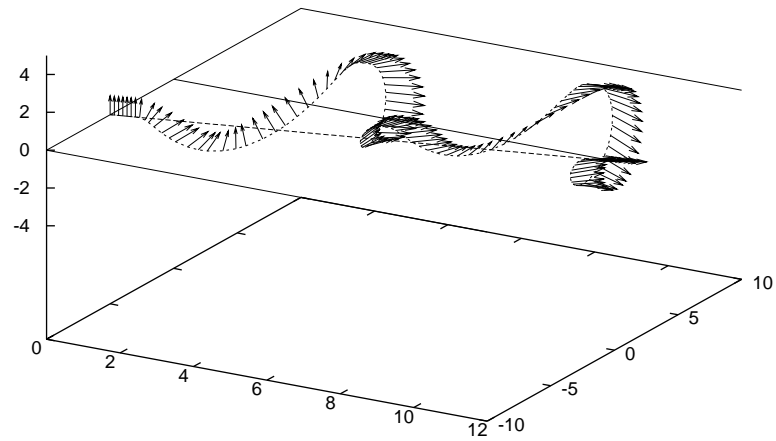
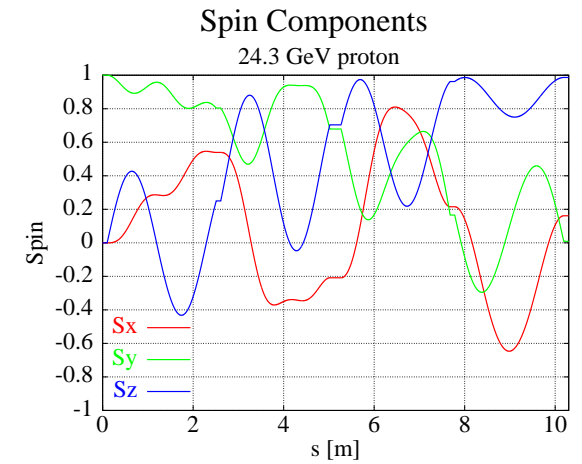
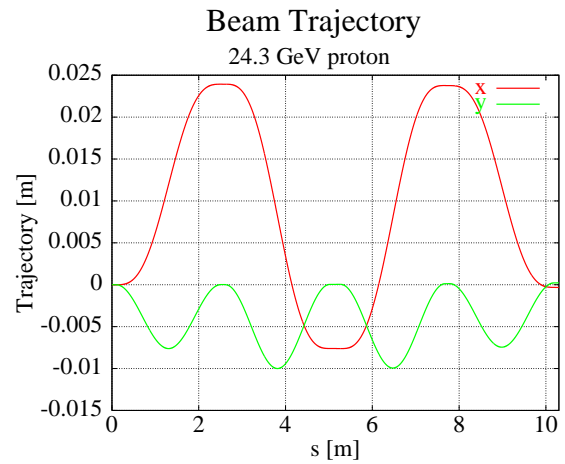
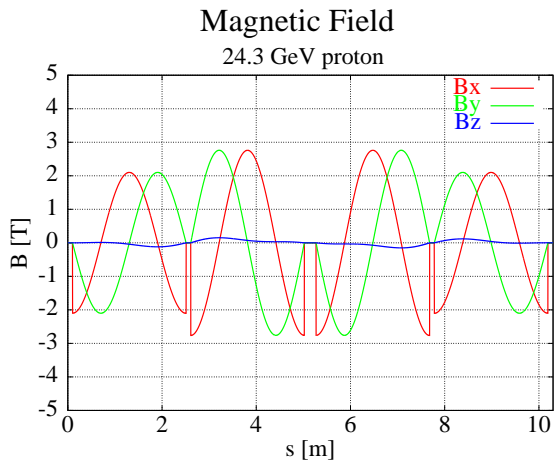


# Snake Resonances

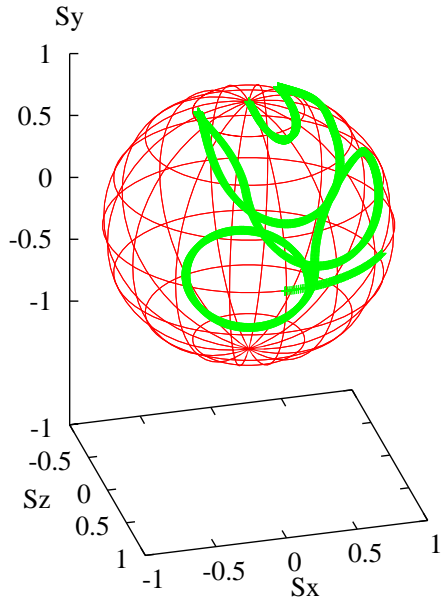
$\epsilon_{\text{int}} = 0.5$ ,  $\epsilon_{\text{imp}} = 0.05$ , 2 Snakes, spin tune = 0.5



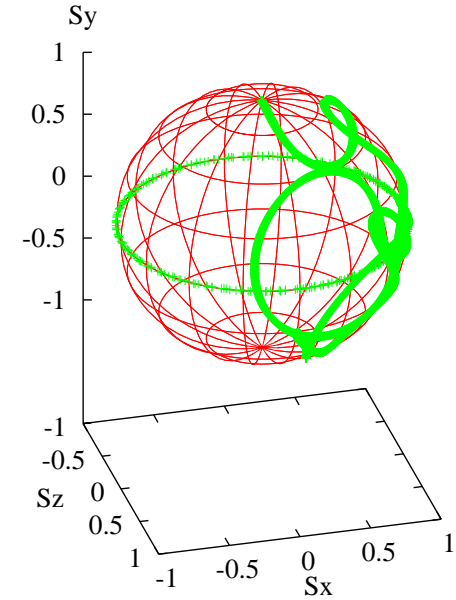
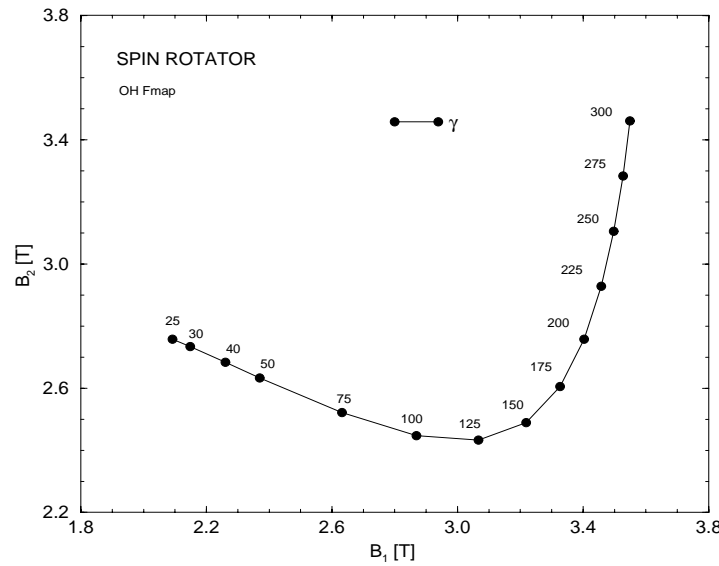
# Helical Spin Rotators



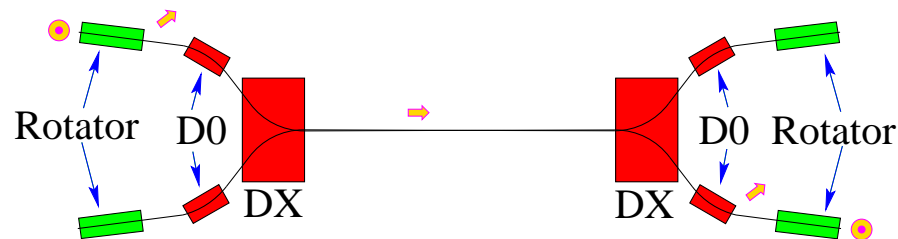
# Compensation for D0-DX Bends



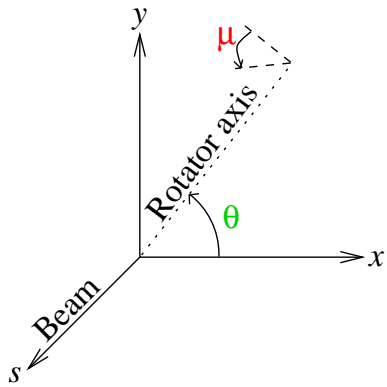
$E = 25 \text{ GeV}$



$E = 250 \text{ GeV}$

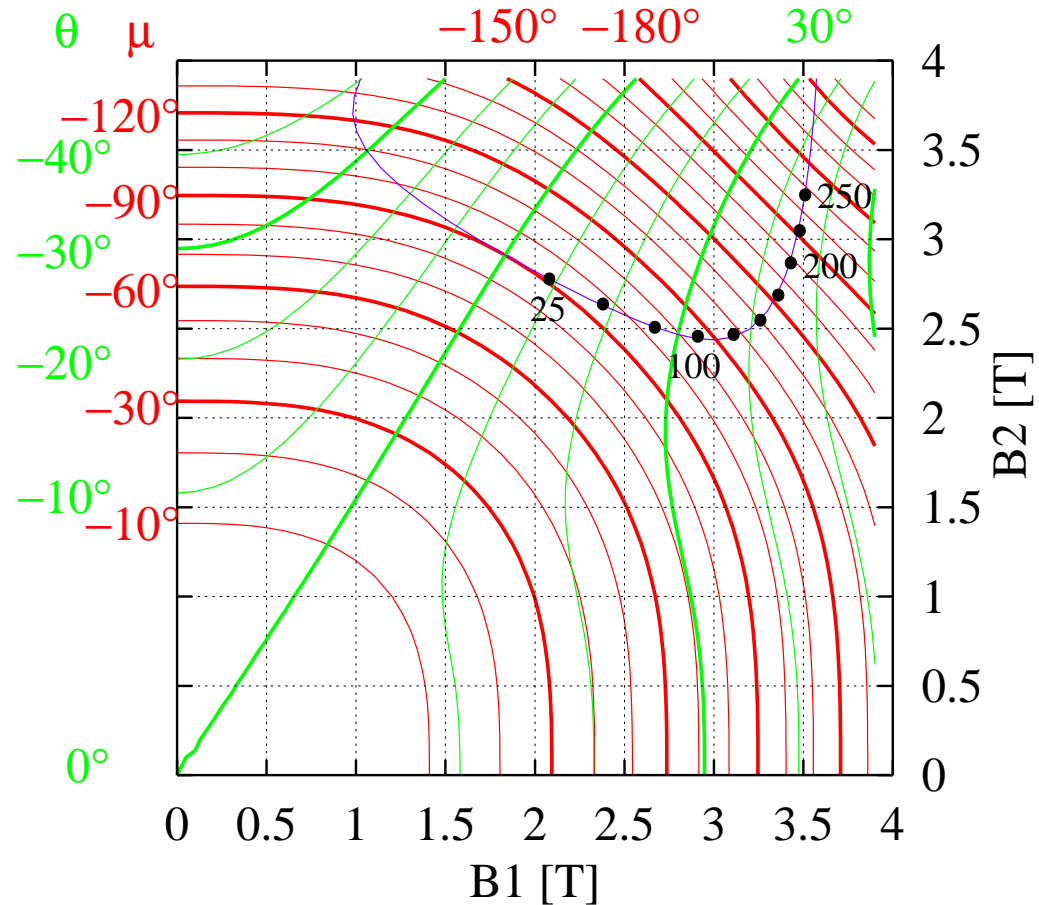






The rotation axis of the spin rotator is in the  $x$ - $y$  plane at an angle  $\theta$  from the vertical. The spin is rotated by the angle  $\mu$  around the rotation axis.

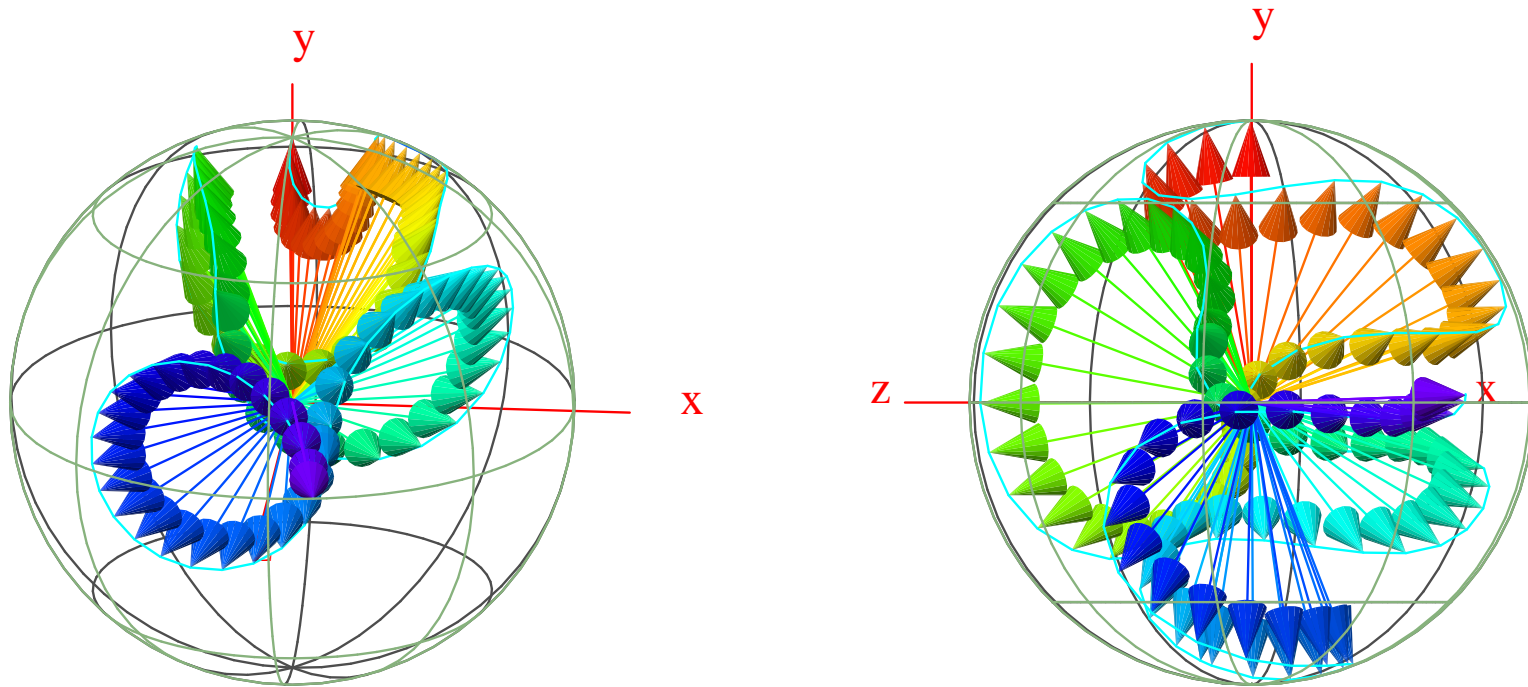
Rotation Angles for a Helical Spin Rotator



Note: Purple contour for rotation into horizontal plane. Black dots show settings for RHIC energies in increments of 25 GeV from 25 to 250 GeV.



# Rotator Spin Precession

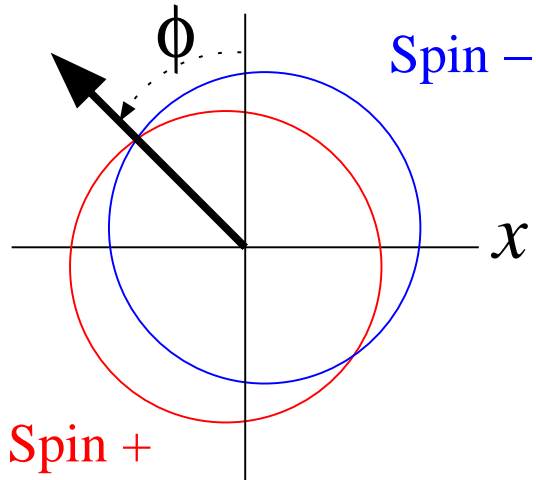


Rotator's spin vector at injection energy

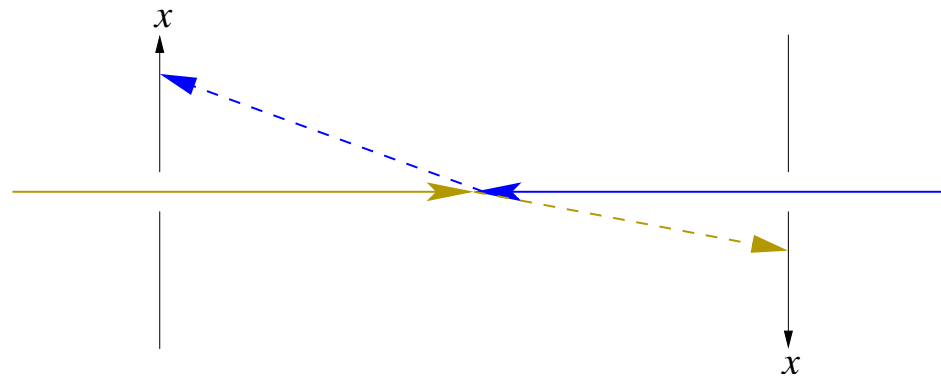
Rotator's spin vector at 250 GeV



# Orientation of PHENIX Polarimeters



"Left-Right" Asymmetry  
(Tilted at  $45^\circ$ )



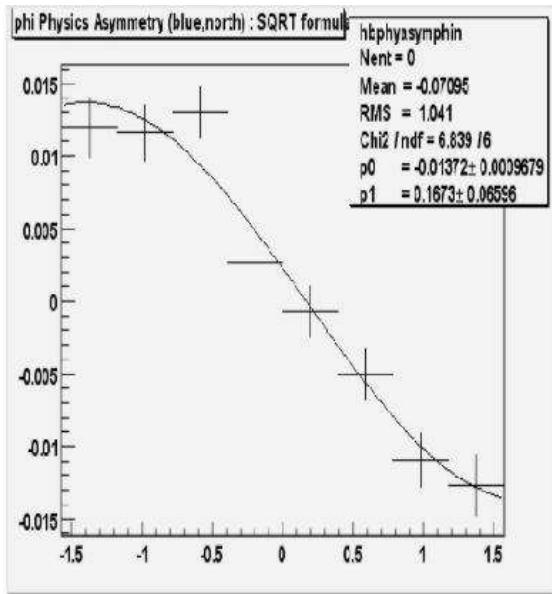
Schematic layout of PHENIX polarimeters  
Yellow from left. Blue from right.

The PHENIX Local Polarimeter measures an asymmetry in small angle scattered neutrons which is proportional to transverse polarization.

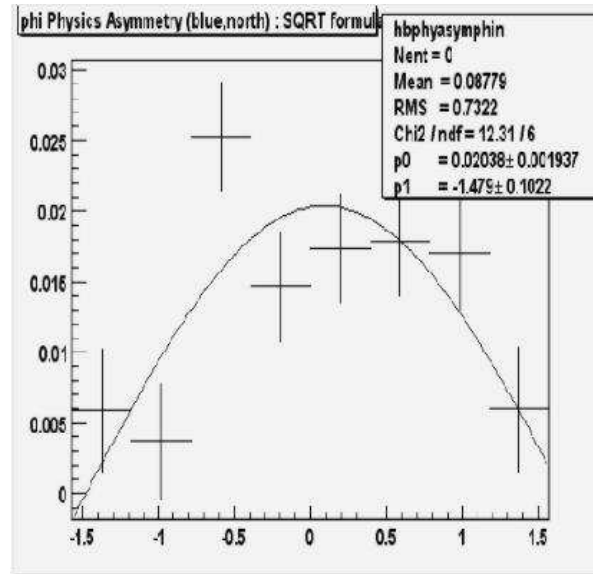
$$A_{LR} = \frac{\sqrt{L^+R^-} - \sqrt{L^-R^+}}{\sqrt{L^+R^-} + \sqrt{L^-R^+}} \propto P_y$$



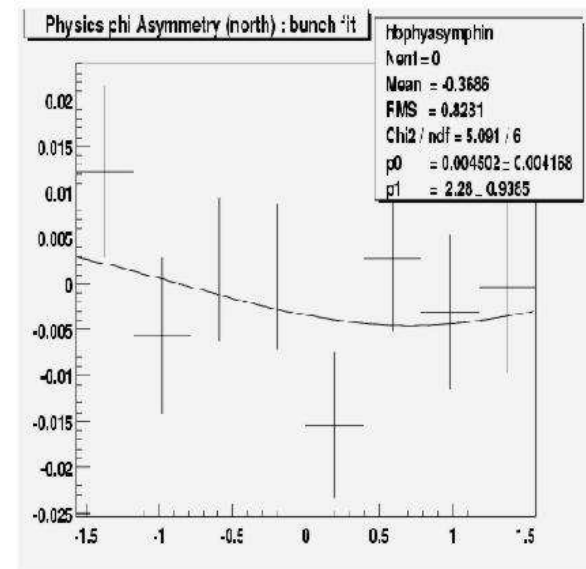
# Tale of the Blue Ring



Vertical polarization  
with rotators off.  
Spin is down.



Rotators on  
Spin is radially inwards!  
OOPS!



Reverse all rotator  
power supplies and try  
again.  
YES!



# Properties of synchrotron radiation

- Radiated power:

$$P_\gamma = \frac{2}{3} r_e m c^3 \frac{\gamma^4 \beta^4}{\rho^2}, \quad r_e = \frac{e^2}{4\pi\epsilon_0 m c^2}.$$

Radiation in forward direction with opening angle  $\propto \gamma^{-1}$

- Energy loss per turn:

$$U_\gamma = \oint \frac{P_\gamma}{c} ds$$

- Critical energy: half the power is radiated by photons less than the critical energy, and the other half, above.

$$u_c = \hbar\omega_c = \frac{3\hbar c}{2\rho} \gamma^3$$



- Number of photons per second:

$$N_\gamma = \int_0^{U_{\max}} n_\gamma(u_\gamma) du_\gamma = \frac{5}{2\sqrt{3}} \frac{\alpha c}{\rho} \gamma$$

here:  $\alpha = 1/137$ )

- Number of photons per radian:

$$N_r = \frac{5\alpha}{2\sqrt{3}} \gamma$$

- Average photon energy and 2<sup>nd</sup> moment:

$$\langle u_\gamma \rangle = \frac{1}{N_\gamma} \int_0^{U_{\max}} u n_\gamma(u) du = \frac{8}{15\sqrt{3}} u_c \simeq 0.32 u_c$$

$$\langle u_\gamma^2 \rangle = \frac{1}{N_\gamma} \int_0^{U_{\max}} u^2 n_\gamma(u) du = \frac{11}{27\sqrt{3}} u_c^2 \simeq 0.41 u_c^2$$



- Energy spread: 
$$\sigma_u = \sqrt{\frac{C_q}{J_s \rho}} \gamma^2 m c^2$$
 with  $C_q = 3.8 \times 10^{-8}$  m and  $J_s \sim 2 + \mathcal{D}$ .

Ring	Energy [GeV]	$\sigma_u$ [MeV]
CESR	5.5	3
HERAe	27.5	3
LEP	45	30
LEP	60	53
LEP	100	150

Remember: Integer resonances separated by only 440 MeV.

The polarization in LEP dropped down to nothing just above 60 GeV.



# Longitudinal Synchrotron Oscillations

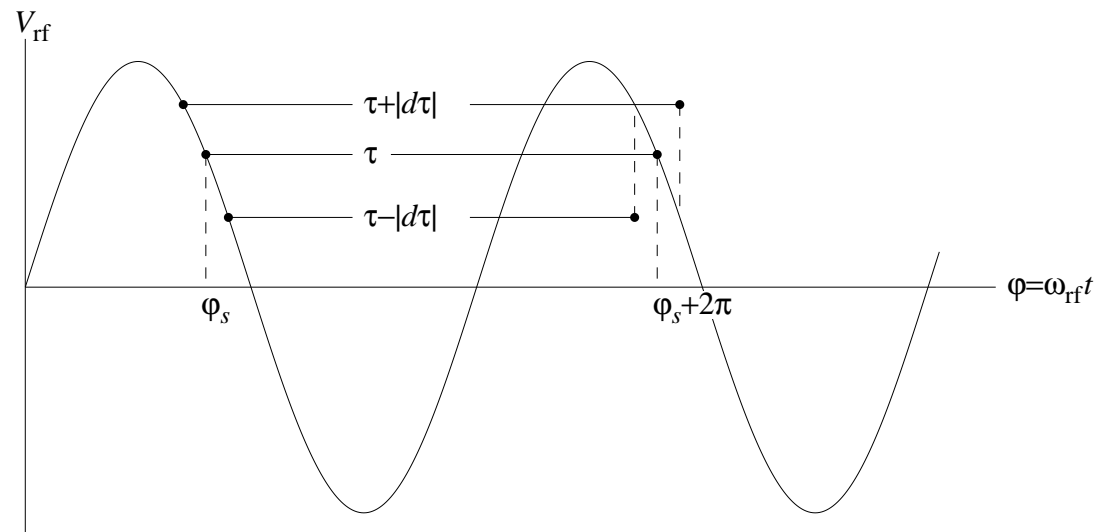
$$\omega_{\text{rf}} = h\omega_{\text{rev}}$$

$$W = -\frac{U - U_s}{\omega_{\text{rf}}}$$

$$\frac{dW}{dt} = \frac{qV}{2\pi h} (\sin \phi_s - \sin \phi)$$

$$\frac{d\phi}{dt} \simeq \frac{\omega_{\text{rf}}^2 \eta_{\text{ph}}}{\beta^2 U_s} W$$

$$\frac{d\omega_{\text{rev}}}{\omega_{\text{rev}}} = \frac{d\beta}{\beta} - \frac{dL}{L} = \eta_{\text{ph}} \frac{dp}{p}$$



$\eta_{\text{ph}} < 0$  above transition energy.

Add in synchrotron oscillations to resonance condition:

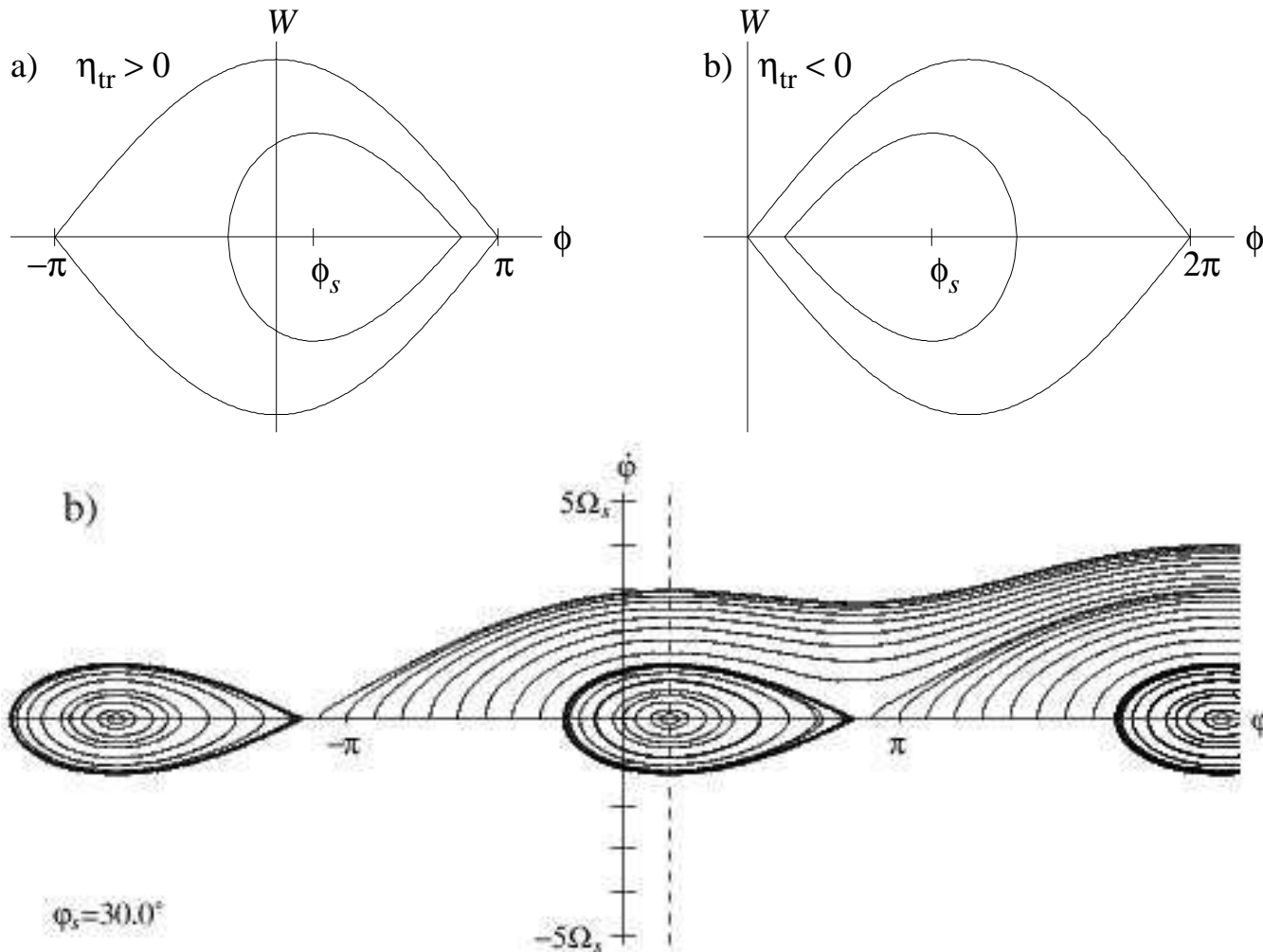
$$\nu_{\text{spin}} = N + N_v Q_v + N_h Q_h + N_{\text{sy}} Q_{\text{sy}}$$





# Longitudinal Phase Space

Canonical coordinate:  $\varphi$  and conjugate momentum:  $W$



# Spin-flip Transition Rates

In a homogenous magnetic field the transition rates are

$$W_{\uparrow\downarrow} = \frac{5\sqrt{3}}{16} \frac{e^2 \gamma^5 \hbar}{4\pi\epsilon_0 m_e c^2 |\rho|^3} \left( 1 + \frac{8}{5\sqrt{3}} \right)$$
$$W_{\downarrow\uparrow} = \frac{5\sqrt{3}}{16} \frac{e^2 \gamma^5 \hbar}{4\pi\epsilon_0 m_e c^2 |\rho|^3} \left( 1 - \frac{8}{5\sqrt{3}} \right).$$

Evaluating the equilibrium polarization have (Sokolov Ternov)

$$P_{\text{ST}} = \frac{W_{\uparrow\downarrow} - W_{\downarrow\uparrow}}{W_{\uparrow\downarrow} + W_{\downarrow\uparrow}} = \frac{8}{5\sqrt{3}} = 0.9238.$$

An unpolarized beam polarizes:

$$P(t) = P_{\text{ST}} [1 - \exp(-t/\tau_{\text{ST}})],$$

where the polarization rate is given by

$$\tau_{\text{ST}}^{-1} = \frac{5\sqrt{3}}{8} \frac{e^2 \gamma^5 \hbar}{4\pi\epsilon_0 m_e^2 c^2} \frac{1}{L} \oint \frac{ds}{|\rho|^3}.$$



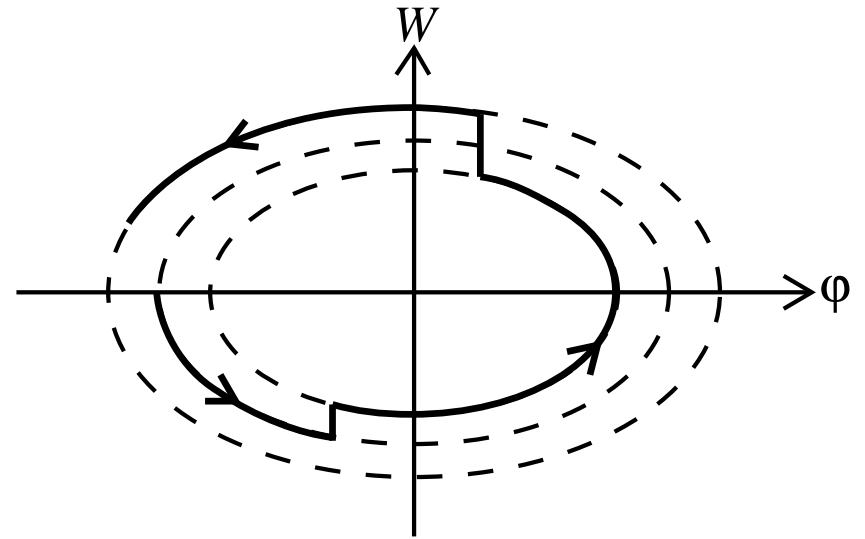
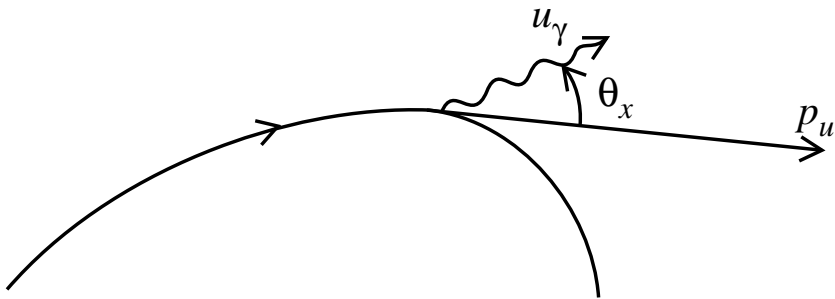
# Typical Sokolov-Ternov Rates

Ring	Particle	Energy [GeV]	$N_\gamma$ [/turn]	$\Delta U$ [loss/turn]	$\tau_{ST}$	$\frac{W_{\uparrow\downarrow}}{f_{rev} N_\gamma}$
CESR	$e^\pm$	5.5	700	-1 MeV	167 min	$1 \times 10^{-13}$
HERAe	$e^\pm$	27.5	3600	-83 MeV	23 min	$1 \times 10^{-12}$
LEP	$e^\pm$	45	5800	-120 MeV	300 min	$2 \times 10^{-13}$
LEP	$e^\pm$	60	7800	-380 MeV	81 min	$8 \times 10^{-13}$
RHIC	p	100	7	-3 meV	$3 \times 10^{14}$ yr	$6 \times 10^{-29}$
RHIC	p	250	18	-0.13 eV	$3 \times 10^{12}$ yr	$2 \times 10^{-27}$
HERAp	p	920	65	-8.5 eV	$1 \times 10^{11}$ yr	$3 \times 10^{-26}$
Tevatron	p	1000	70	-8.5 eV	$2 \times 10^{11}$ yr	$2 \times 10^{-26}$
SSC	p	20000	1400	-0.12 MeV	$7 \times 10^7$ yr	$3 \times 10^{-23}$

Age of the universe  $\sim 13.8 \times 10^9$  yr.



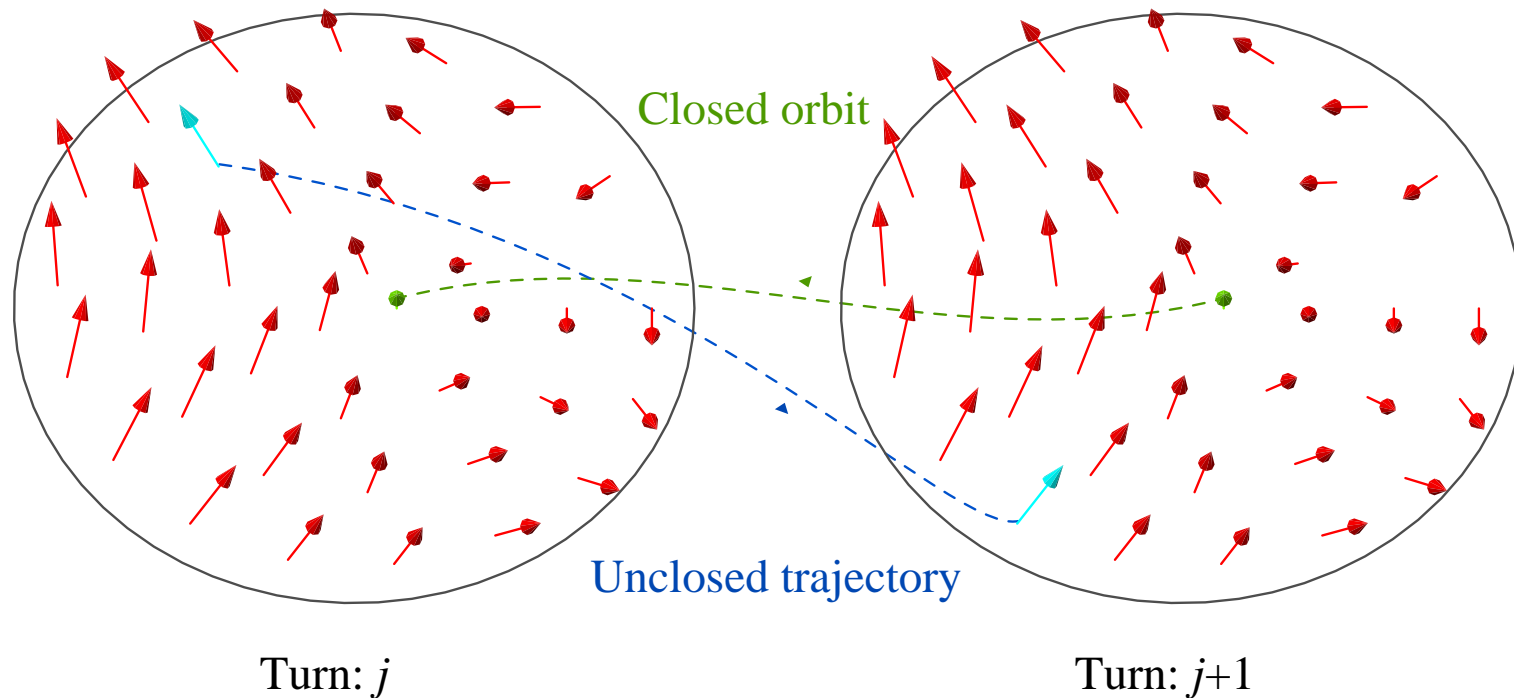
# Quantum Fluctuations



In phase space quantum fluctuations cause instantaneous hops of momentum from one ellipse to another. (Hops in the Action.)



# Invariant Spin Field



- For the closed orbit:  $\vec{n}_0(s) = \vec{n}_0(s + L)$ ,  
with  $\vec{q}_0(s) = \vec{q}_0(s + L)$  and  $\vec{P}_0(s) = \vec{P}_0(s + L)$ .
- For other locations in phase space:  $\vec{n}(\vec{q}, \vec{P}, s) = \vec{n}(\vec{q}, \vec{P}, s + L)$ ,  
even though in general  $q(s + L) \neq q(s)$  and  $P(s + L) \neq P(s)$ .



# Equilibrium with Real Lattice

Derbenev–Kondratenko formula for equilibrium polarization:

$$P_{\text{DK}} = \frac{8}{5\sqrt{3}} \frac{\oint \left\langle \frac{1}{|\rho|^3} \hat{b} \cdot \left( \hat{n} - \frac{\partial \hat{n}}{\partial \delta} \right) \right\rangle_s ds}{\oint \left\langle \frac{1}{|\rho|^3} \left[ 1 - \frac{2}{9} (\hat{n} \cdot \hat{s})^2 + \frac{11}{18} \left( \frac{\partial \hat{n}}{\partial \delta} \right)^2 \right] \right\rangle_s ds}$$

$$\frac{1}{\tau_{\text{DK}}} = \frac{5\sqrt{3}}{8} \frac{r_e \gamma^5 \hbar}{m_e} \frac{1}{L} \oint \left\langle \frac{1}{|\rho|^3} \left[ 1 - \frac{2}{9} (\hat{n} \cdot \hat{s}) + \frac{11}{18} \left( \frac{\partial \hat{n}}{\partial \delta} \right)^2 \right] \right\rangle_s ds$$

averaged over phase space at azimuth  $s$ .

$\delta = \Delta p/p$  is the fractional momentum deviation from design.

$\hat{n}$  is the invariant spin field.

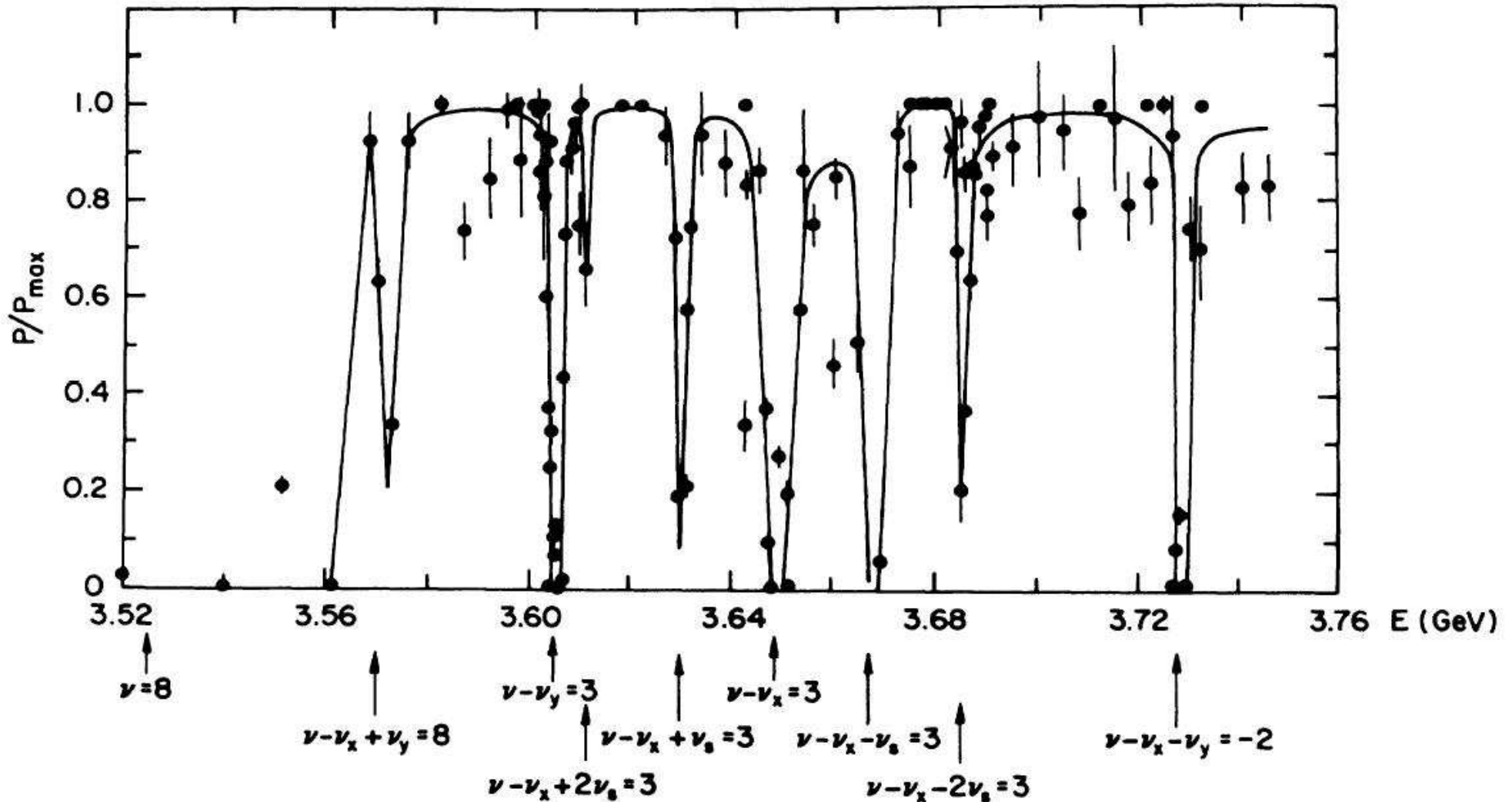
$\hat{b} = \frac{\hat{s} \times \dot{\hat{s}}}{|\dot{\hat{s}}|}$  is the direction of magnetic field if  $\vec{E} = 0$ .

$\rho$  is the cyclotron radius of the trajectory.

$L$  is circumference of synchrotron.



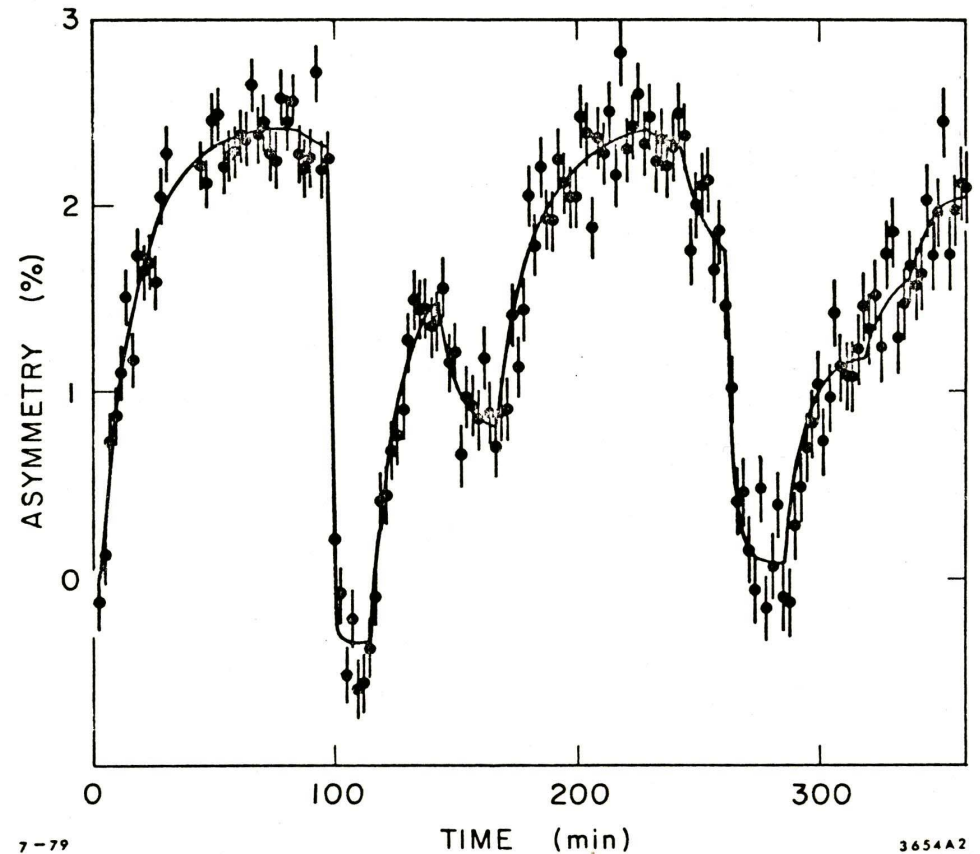
# Spin Resonances of SPEAR



# SPEAR Polarization Time

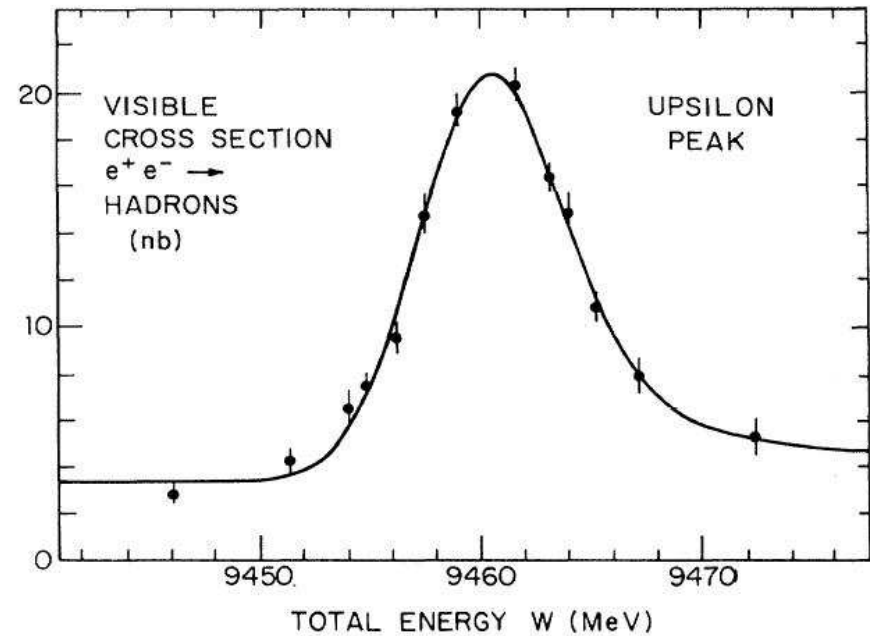
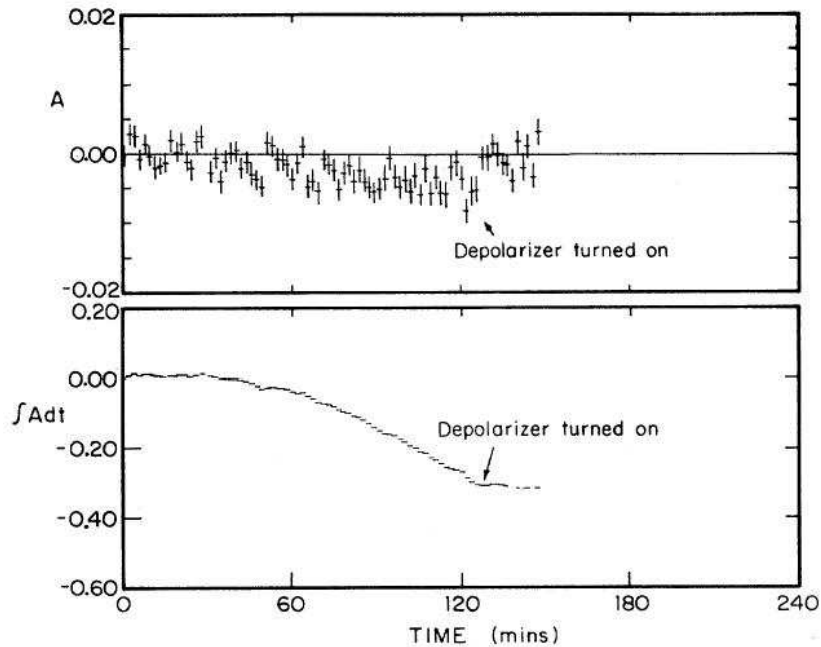
$$\frac{1}{\tau_{\text{dep}}} \approx \frac{5\sqrt{3}}{8} \frac{r_e \gamma^5 \hbar}{m_e} \frac{1}{L} \oint \left\langle \frac{1}{|\rho|^3} \left( \frac{\partial \hat{n}}{\partial \delta} \right)^2 \right\rangle_s ds$$

$$\frac{1}{\tau_{\text{pol}}} \approx \frac{1}{\tau_{\text{ST}}} - \frac{1}{\tau_{\text{dep}}}$$





# Precision Mass Measurements

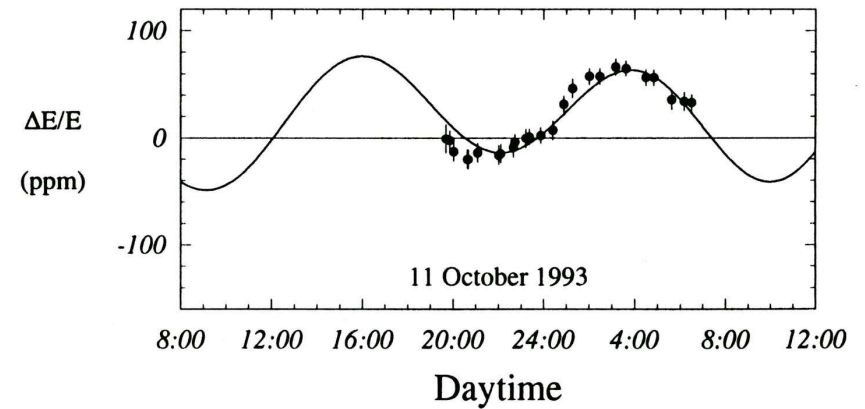
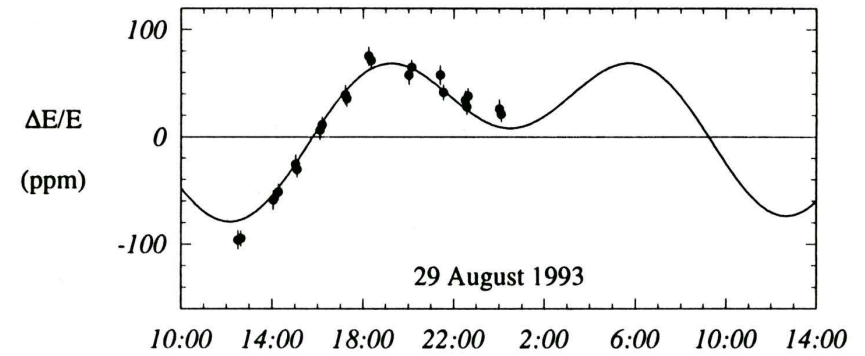
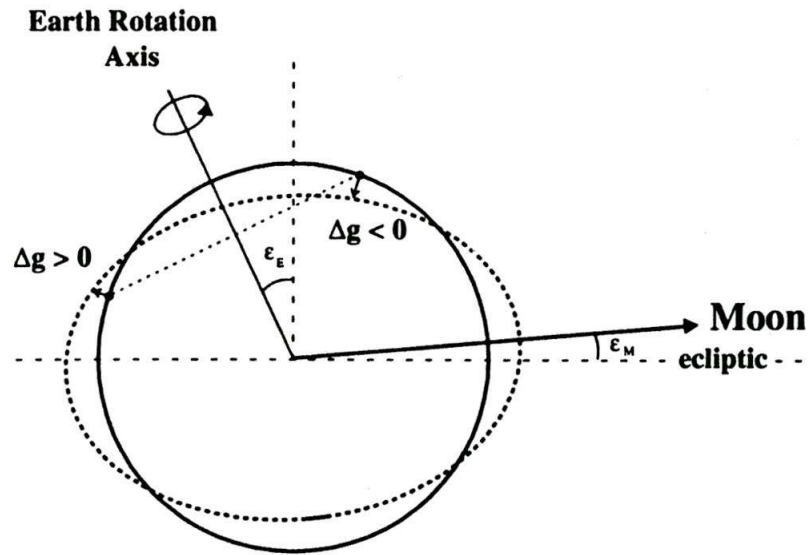


As an example from CESR (CUSB):  $M_\Upsilon = 9459.97 \pm 0.11 \pm 0.07$  MeV

[Phys Rev D29, 2483 (1984)].



# Tidal Effects at LEP



From Angelika Drees' Thesis



# Some References (by no means all!)

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